

## Experimental study of accretion processes in X-ray binary stars using an external magnetic field

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**hzdr**



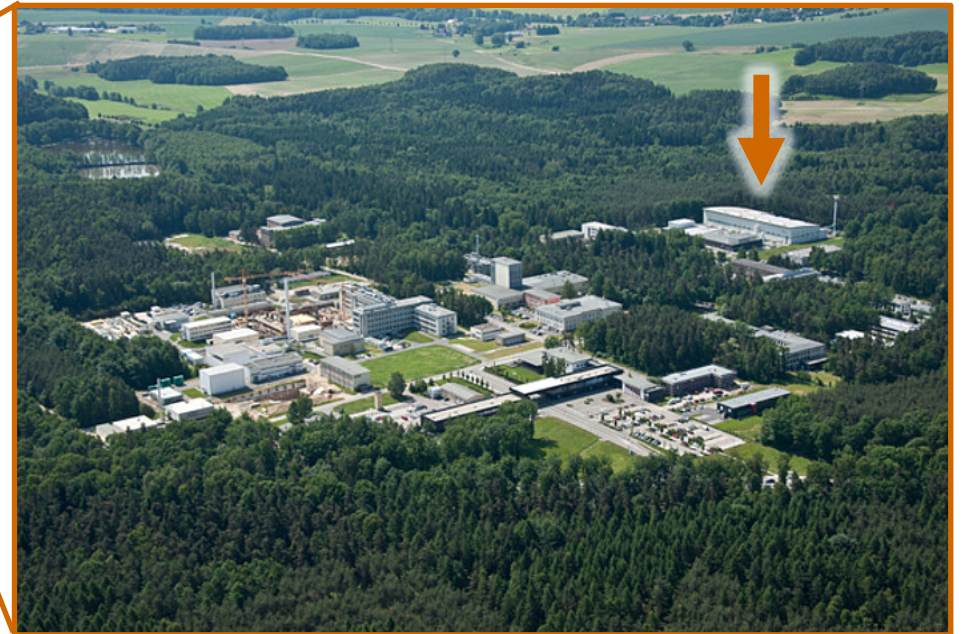
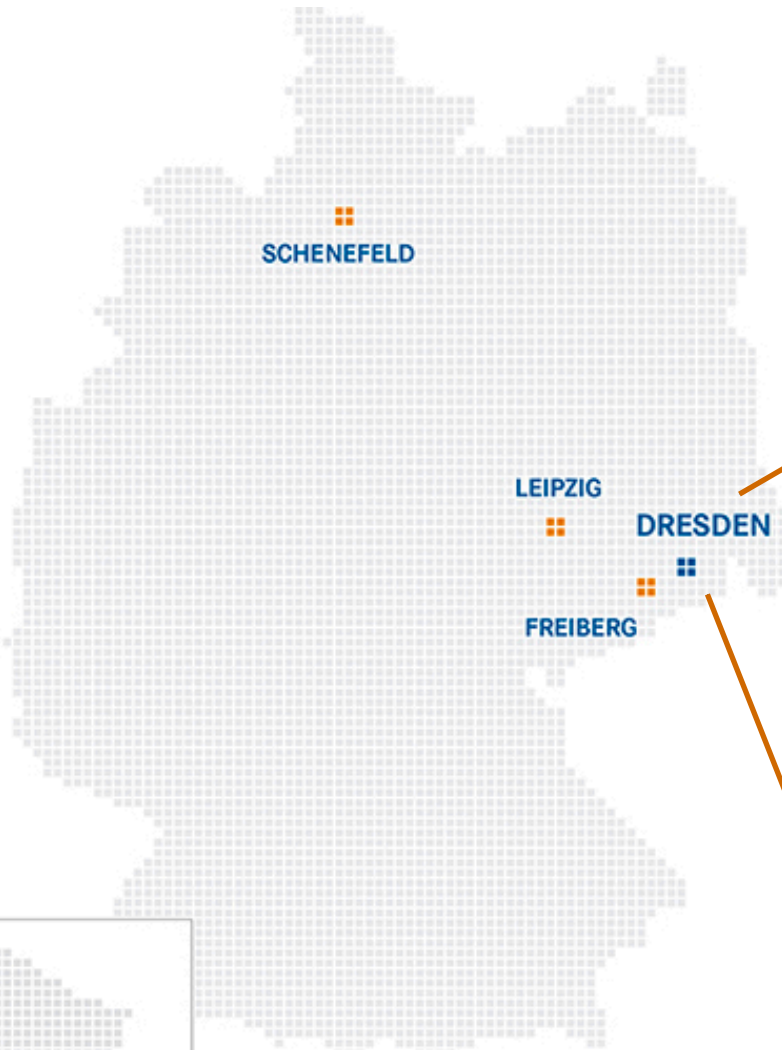
HELMHOLTZ  
ZENTRUM DRESDEN  
ROSSENDORF



Hi I am Florian, but you can call me Flo!

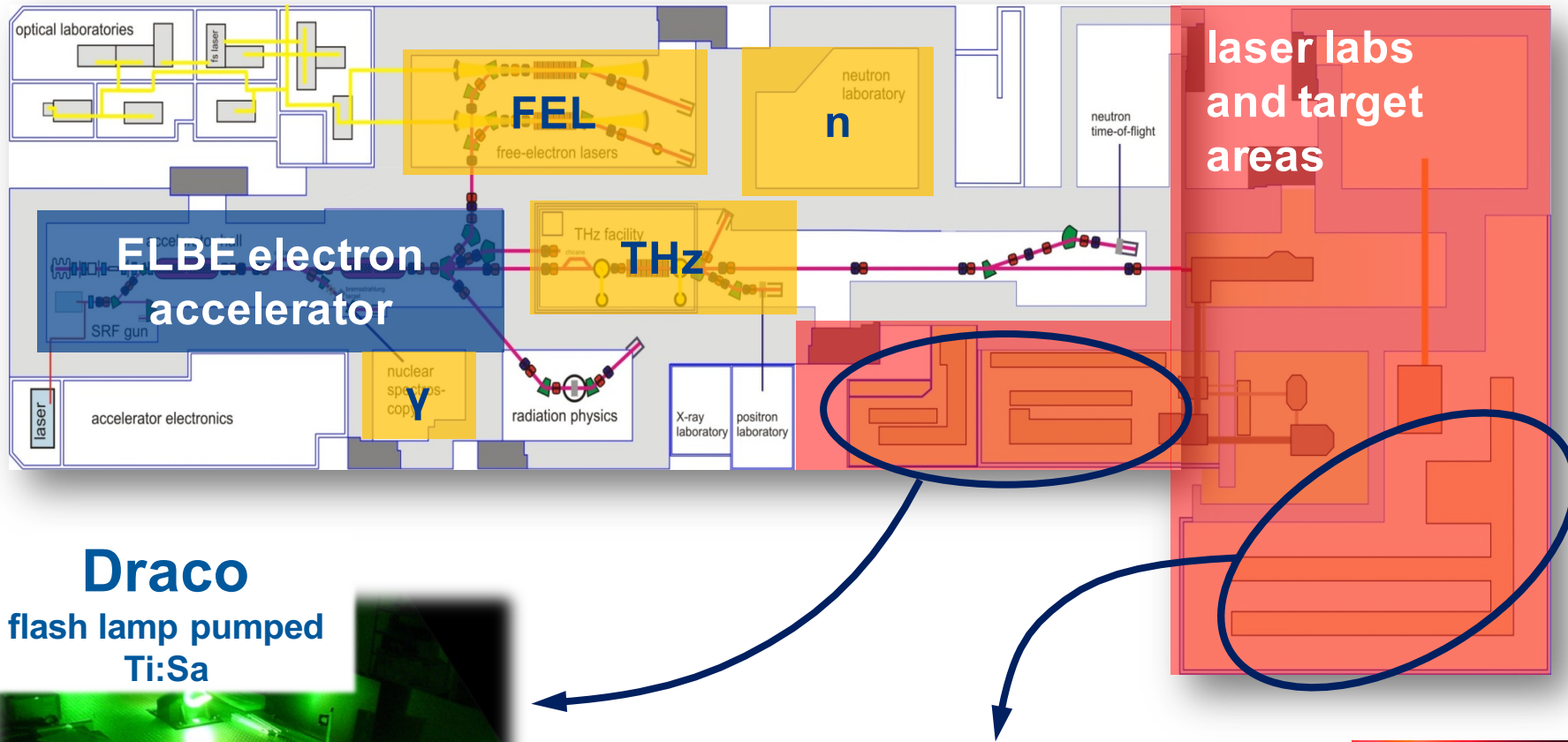
**HZDR**

**HELMHOLTZ  
ZENTRUM DRESDEN  
ROSSENDORF**

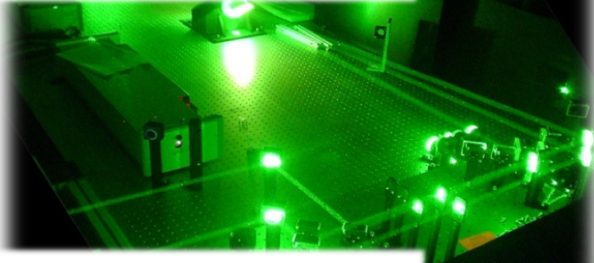


**HZDR**

# The laser particle acceleration division @ HZDR



**Draco**  
flash lamp pumped  
Ti:Sa



Dual beam:  
4J in 30 fs  
+ 30 J in 30 fs

**PENELOPE**  
diode pumped  
Yb:CaF

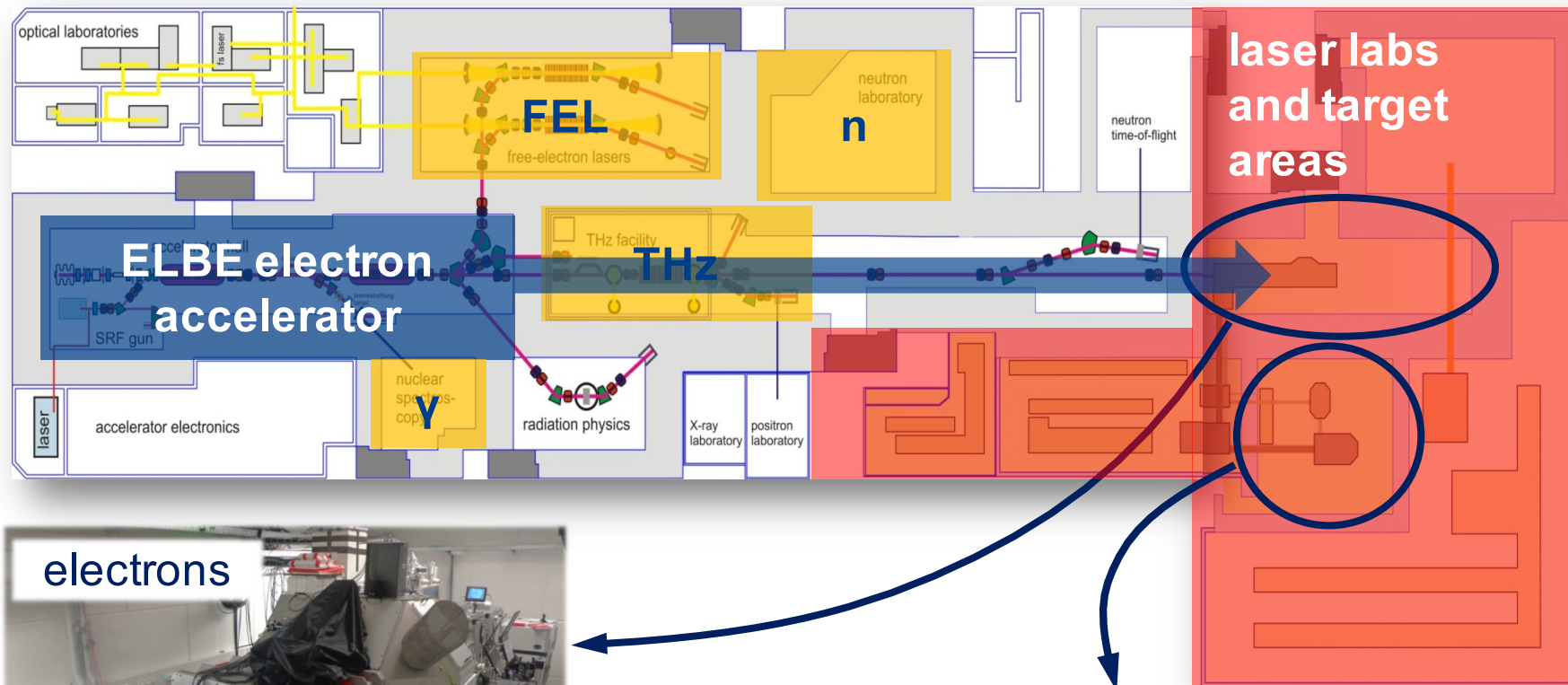


150 J in 150 fs

**Petawatt Laser Power**  
**1 Hz repetition rate**



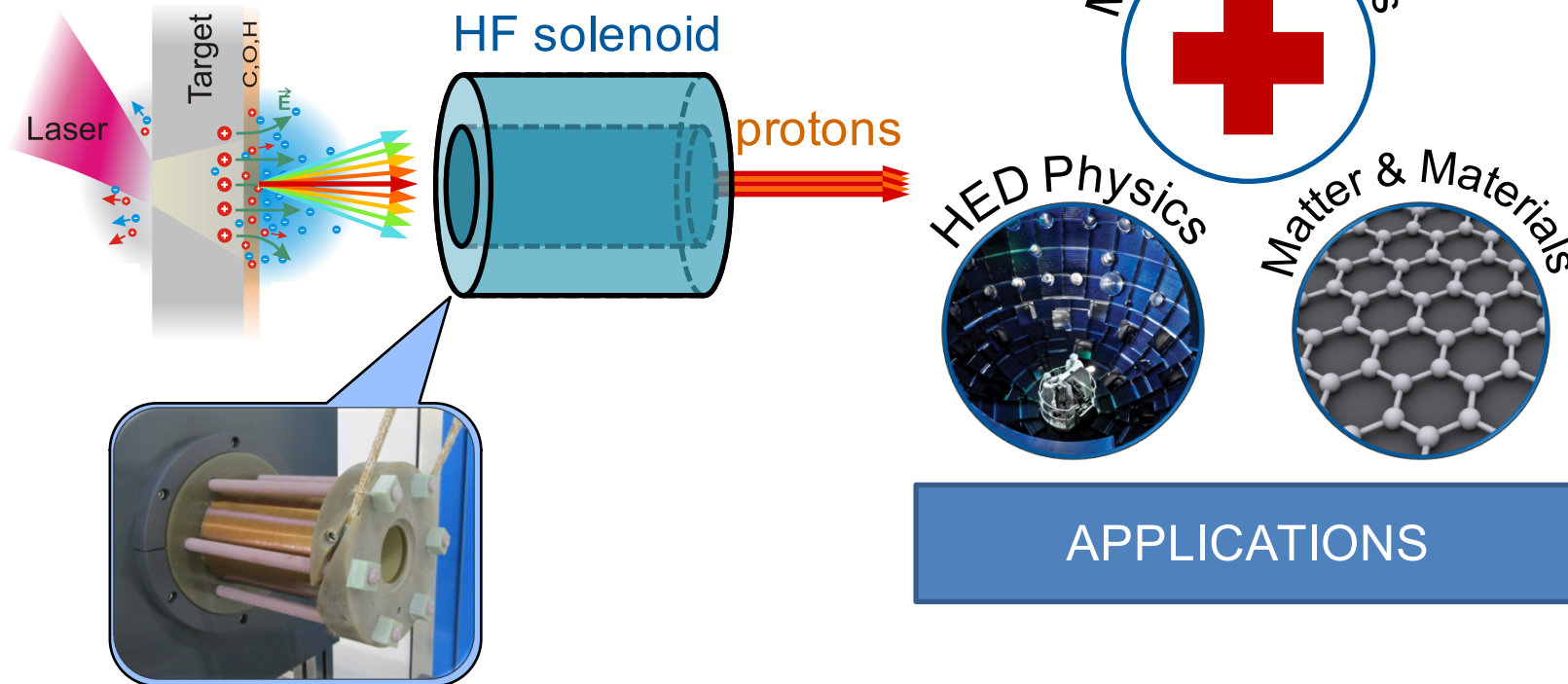
# The laser particle acceleration division @ HZDR





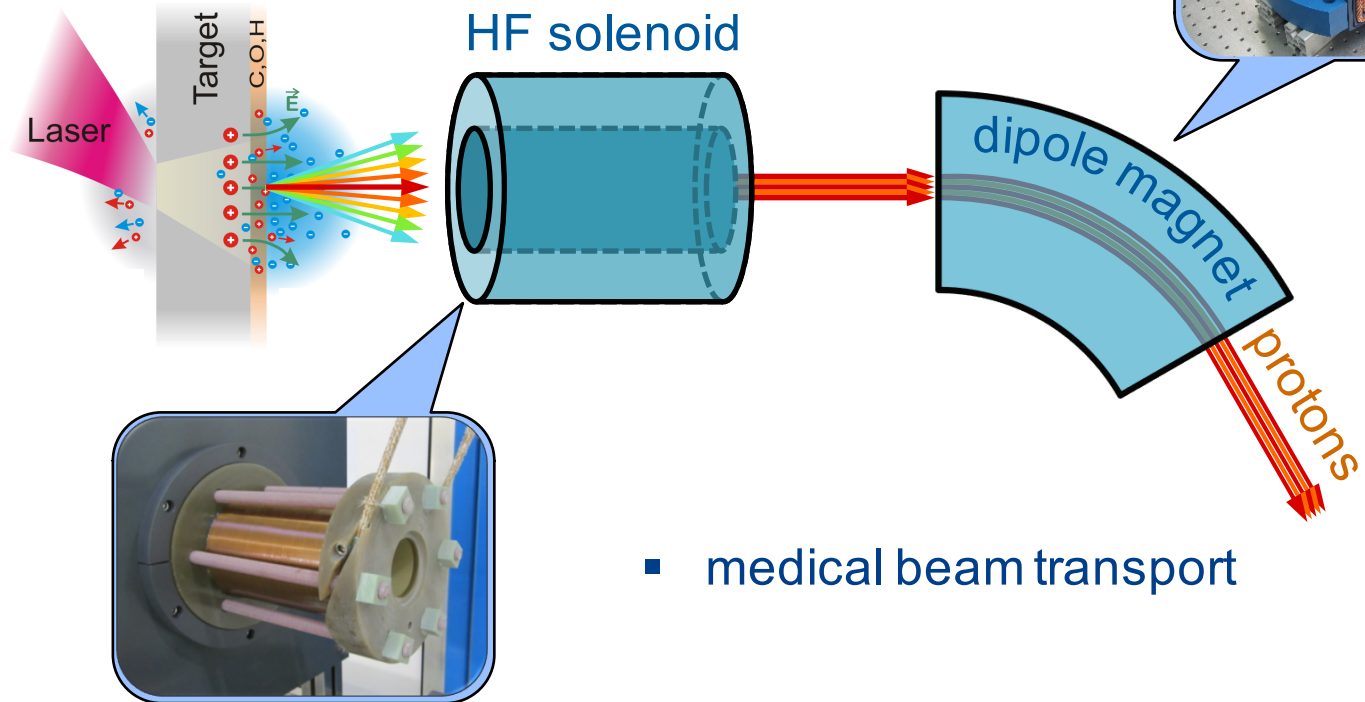
# What the heck is this guy doing here?

## Pulsed high-field magnet development



# What the heck is this guy doing here?

## Pulsed high-field magnet development



- medical beam transport

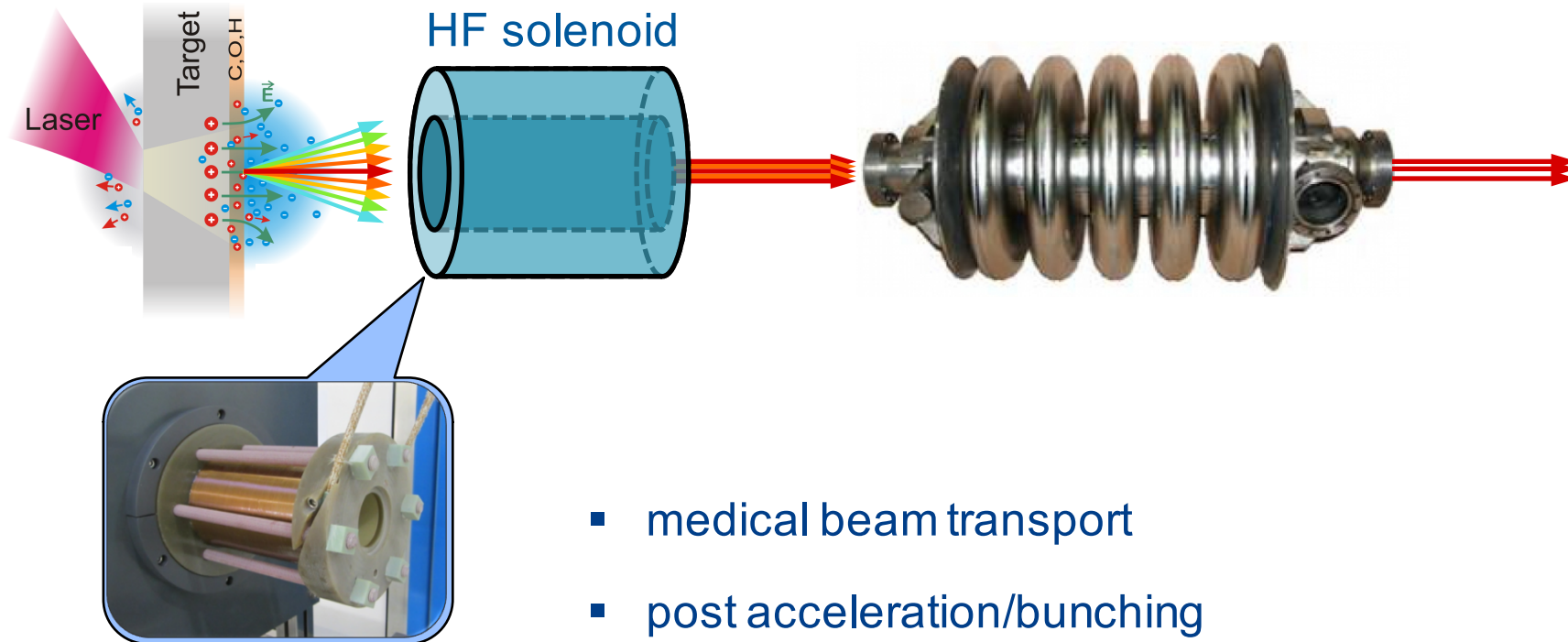
U. Masood et al., APB 117, 41-52 (2014)





# What the heck is this guy doing here?

## Pulsed high-field magnet development



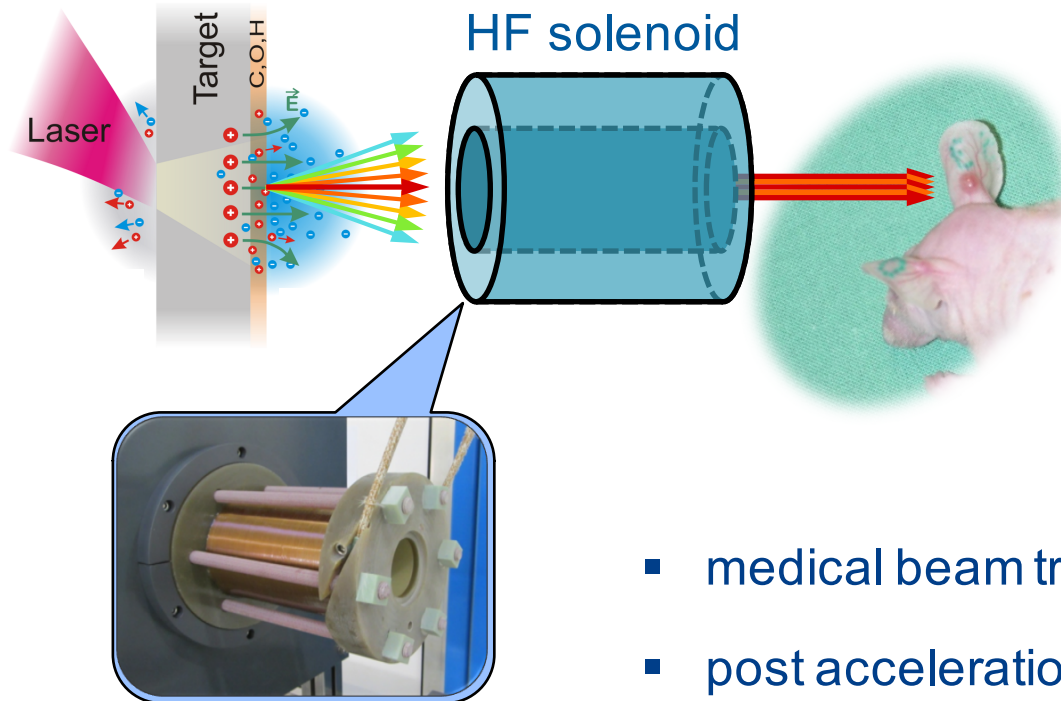
- medical beam transport
- post acceleration/bunching

S. Busold et al., PRSTAB 17, 031302 (2014)  
S. Busold et al., NIM-A 740, 94-98 (2014)



# What the heck is this guy doing here?

## Pulsed high-field magnet development



- medical beam transport
- post acceleration/bunching
- radiobiology

K. Brüchner et al., RadiatOncol, 9:57 (2014)



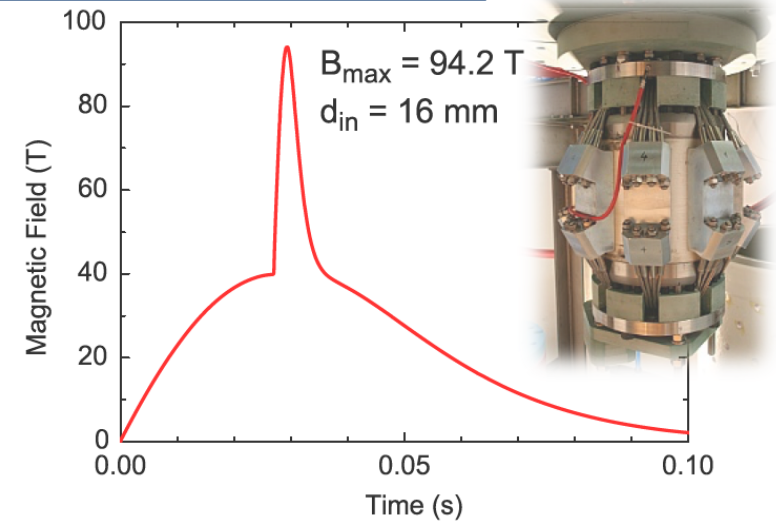
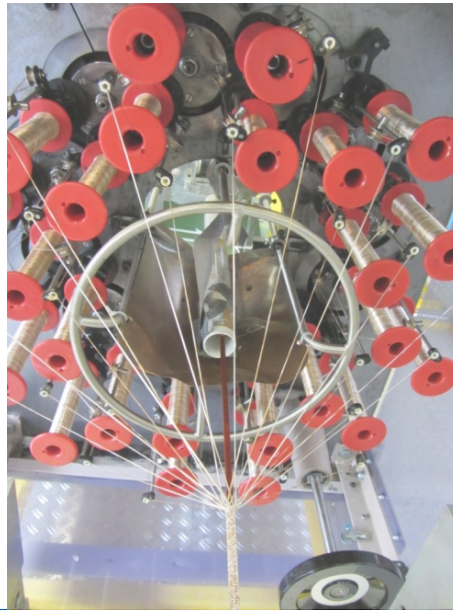


# Dresden High Magnetic Field Laboratory

**HLD.**

Variety of promising application scenarios in  
lab astrophysics...

... and always opportunities to collaborate!



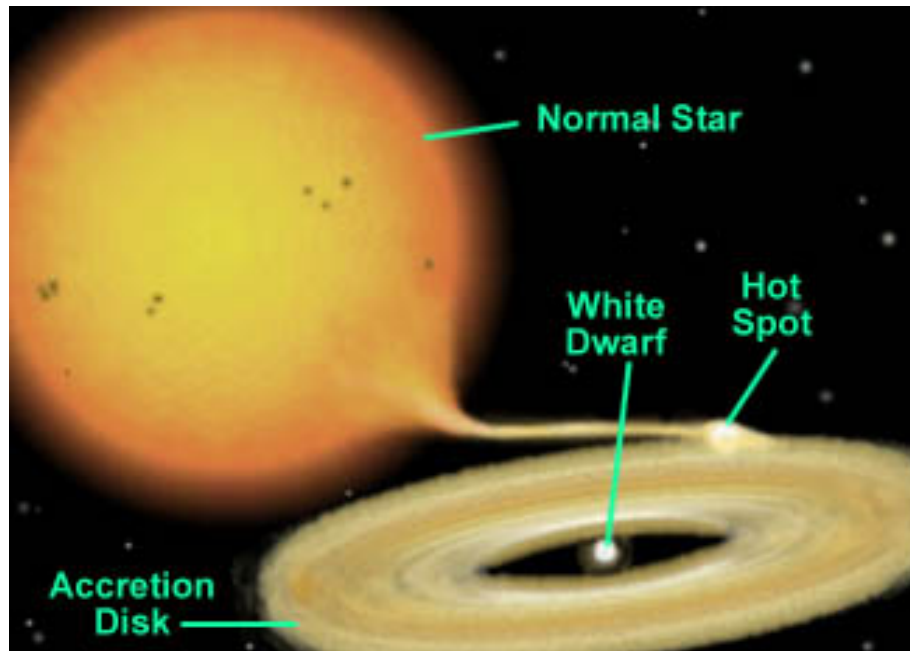
# Binary star systems

## What is a cataclysmic variable star?

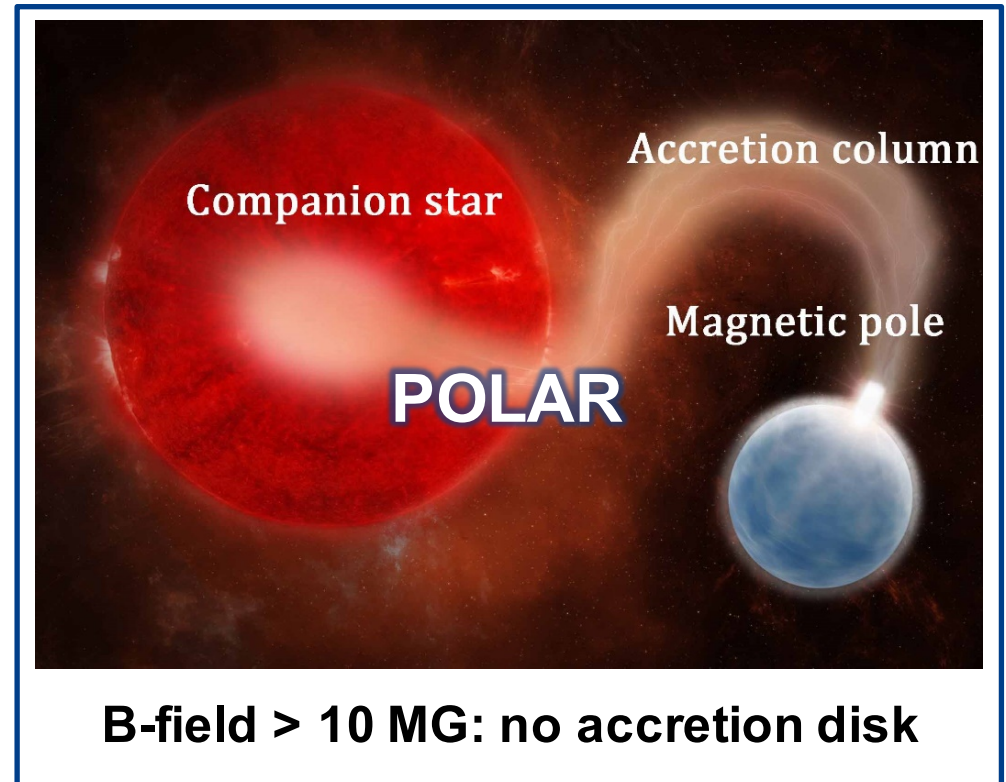
=

A binary system composed of a white dwarf accreting matter from a sun-like star

### Two types of CVs:



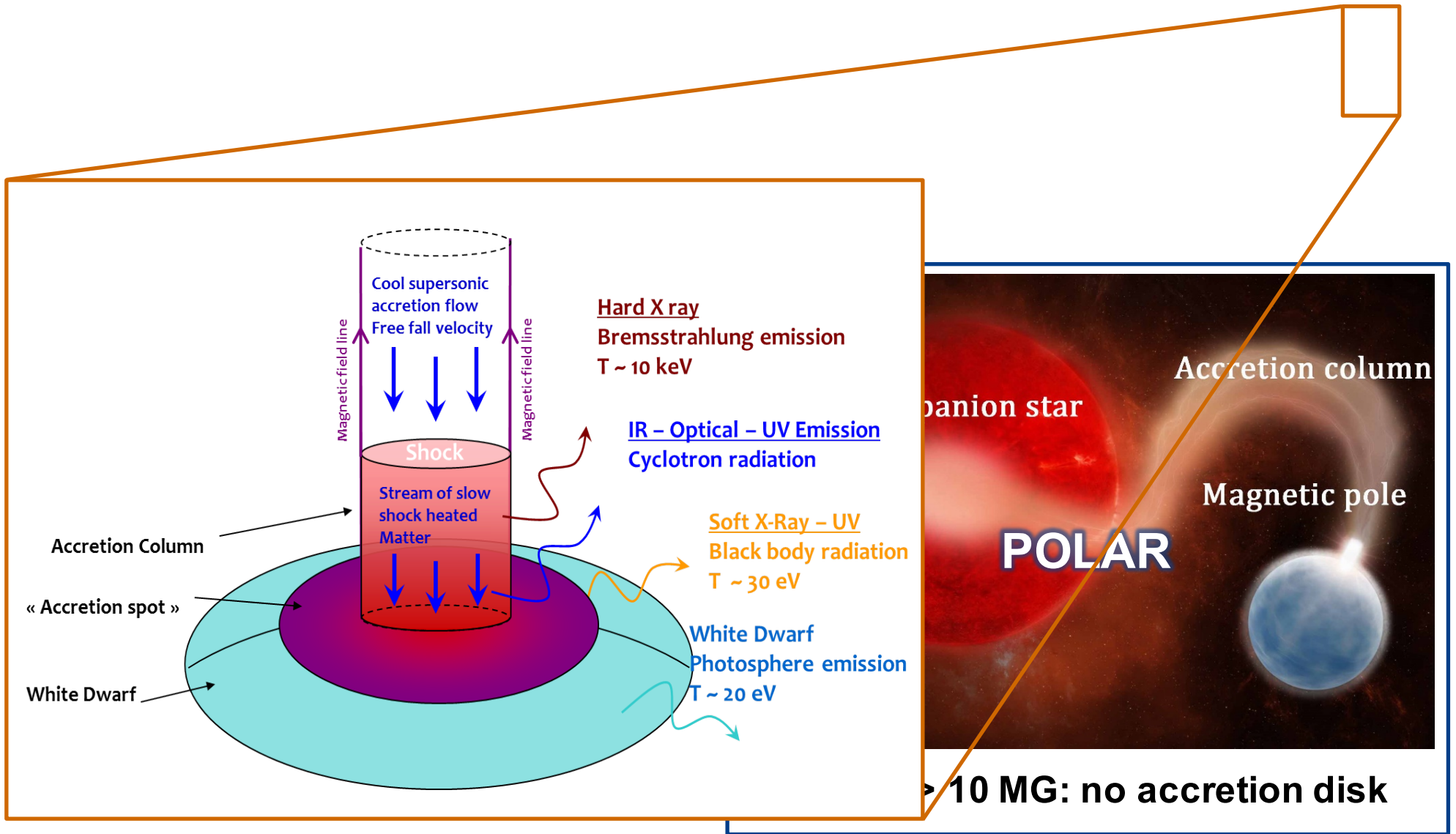
**B-field < 1 MG: accretion disk**



**B-field > 10 MG: no accretion disk**

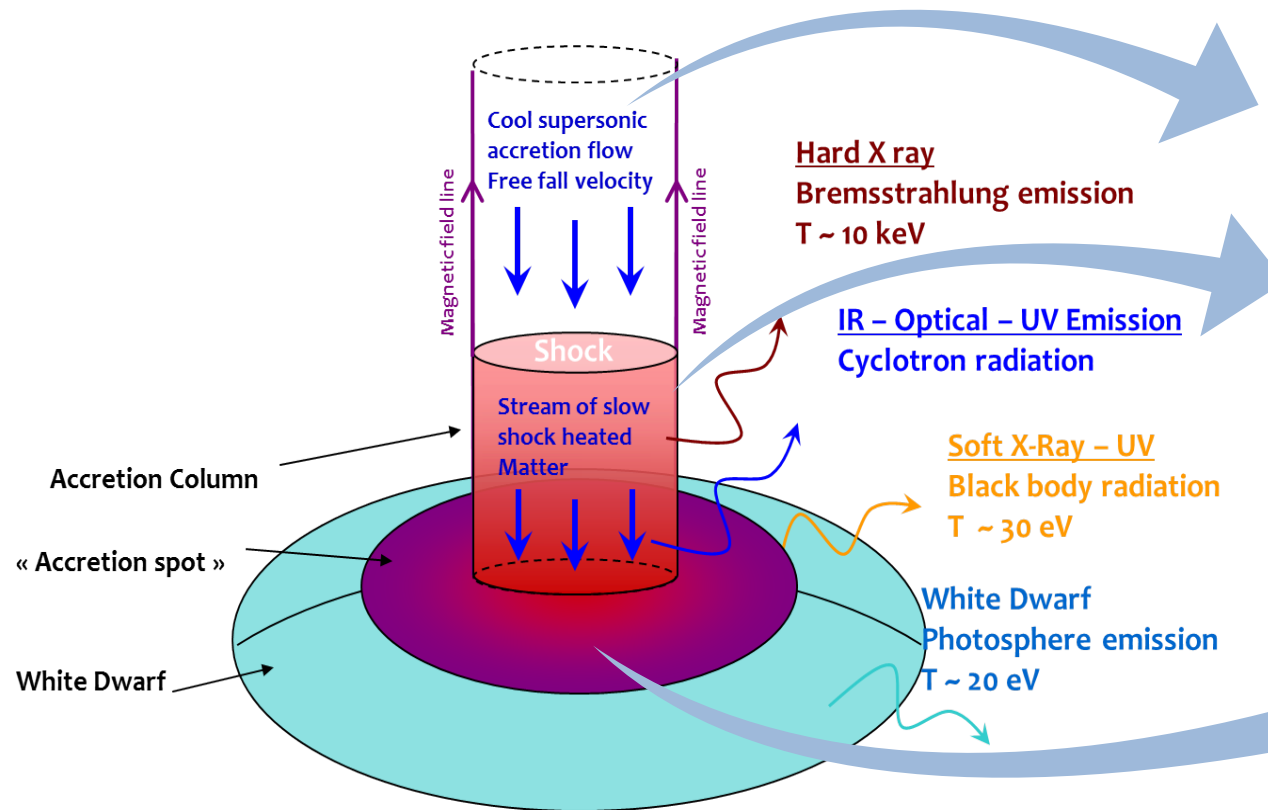


# Binary star systems



# Binary star systems in the lab?

## Astrophysical scenario:



## Experiment:

Reproduce a plasma flow:  
laser-target interaction

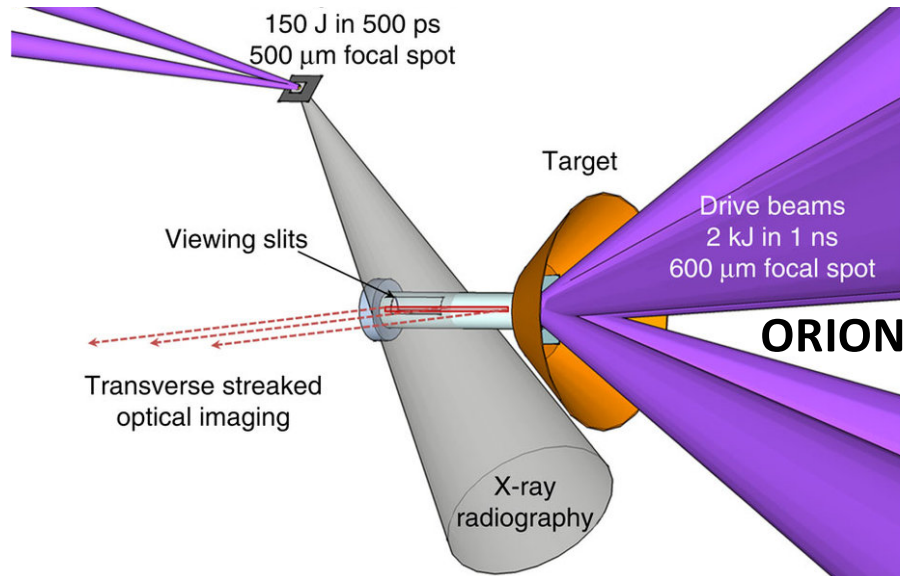
Collimate the plasma flow:  
???

Simulate WD photosphere:  
obstacle

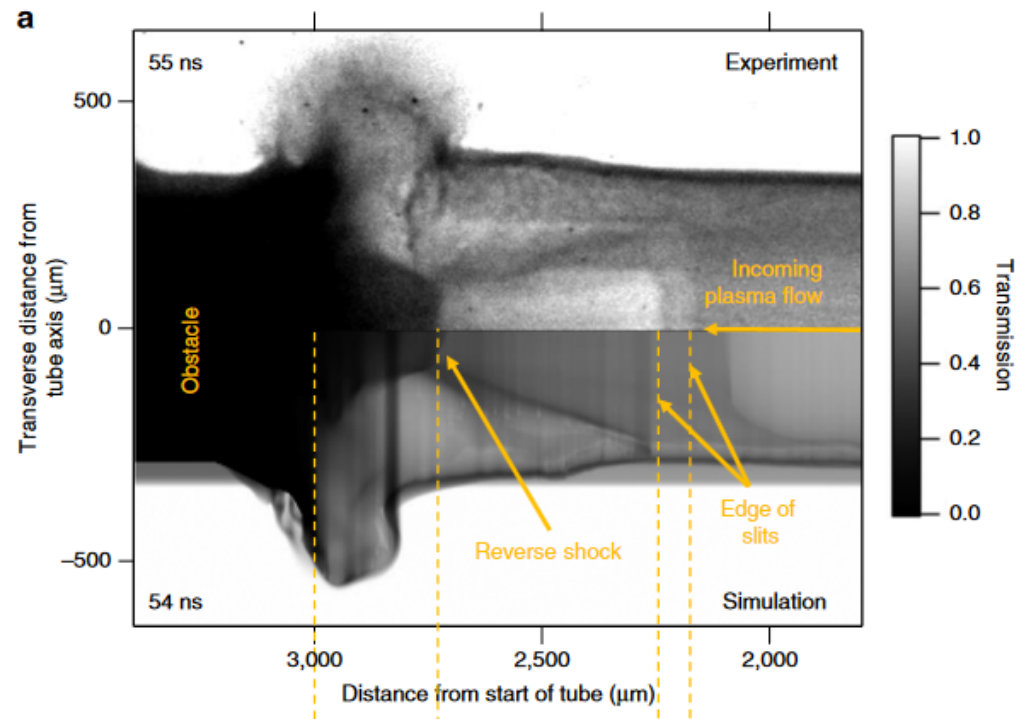


# Plasma collimation via tube

ORION + LULI2000 (2012, 2014)



J. Cross et al., Nat Commun, 7:11899 (2016)



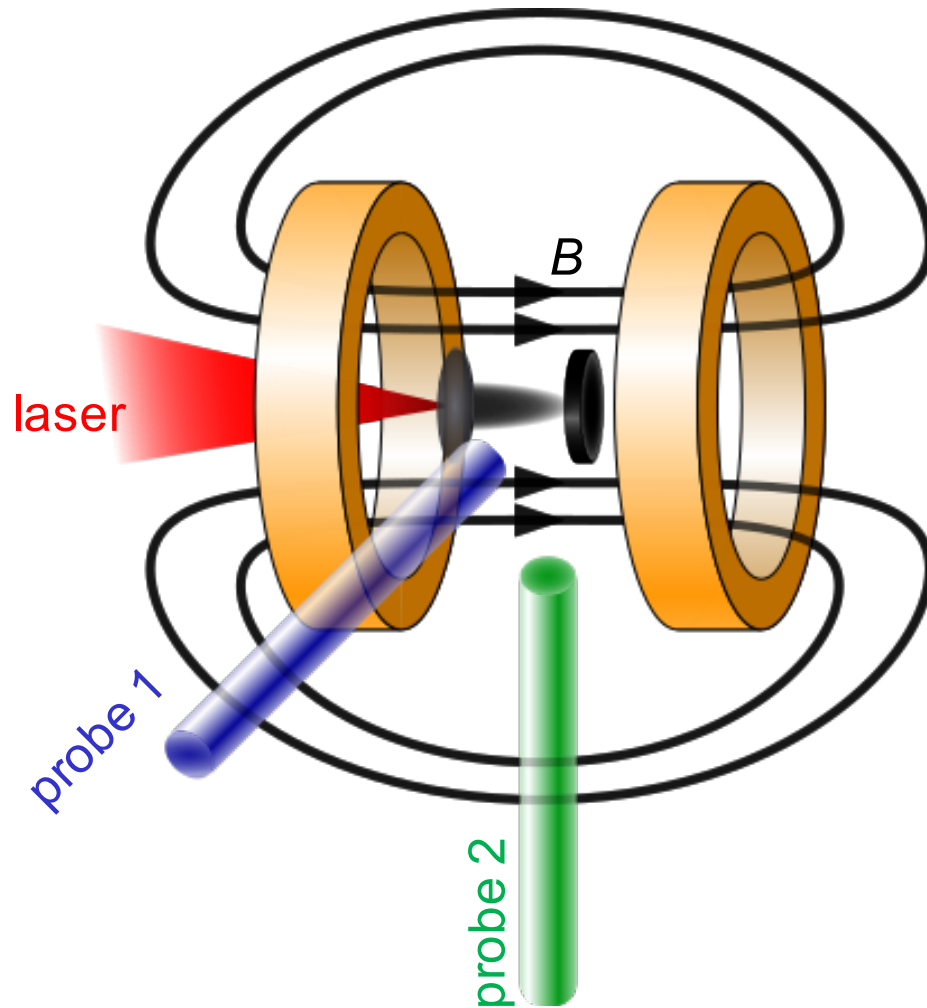
- simulations and experiment in good agreement
- **Tube strongly influences plasma dynamics**

Can we accomplish the same with a magnetic field and still study the jet expansion and shock physics?

cf. B. Albertazzi, Science, 346, 325-328 (2014)



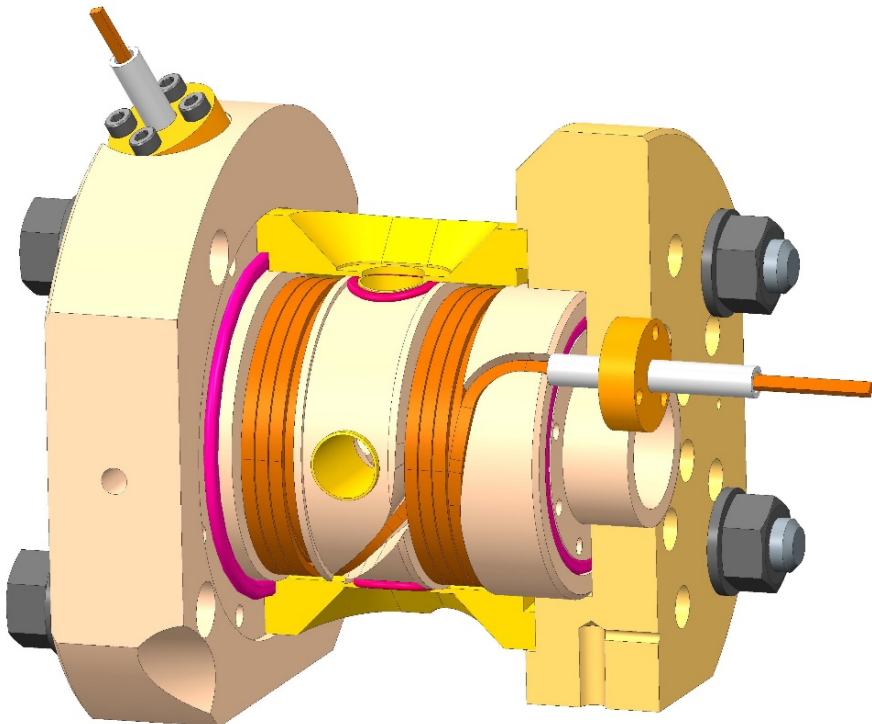
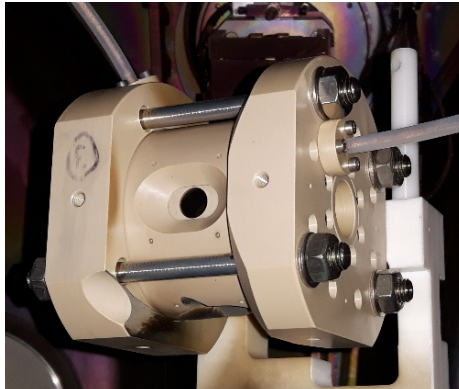
# Experimental setup



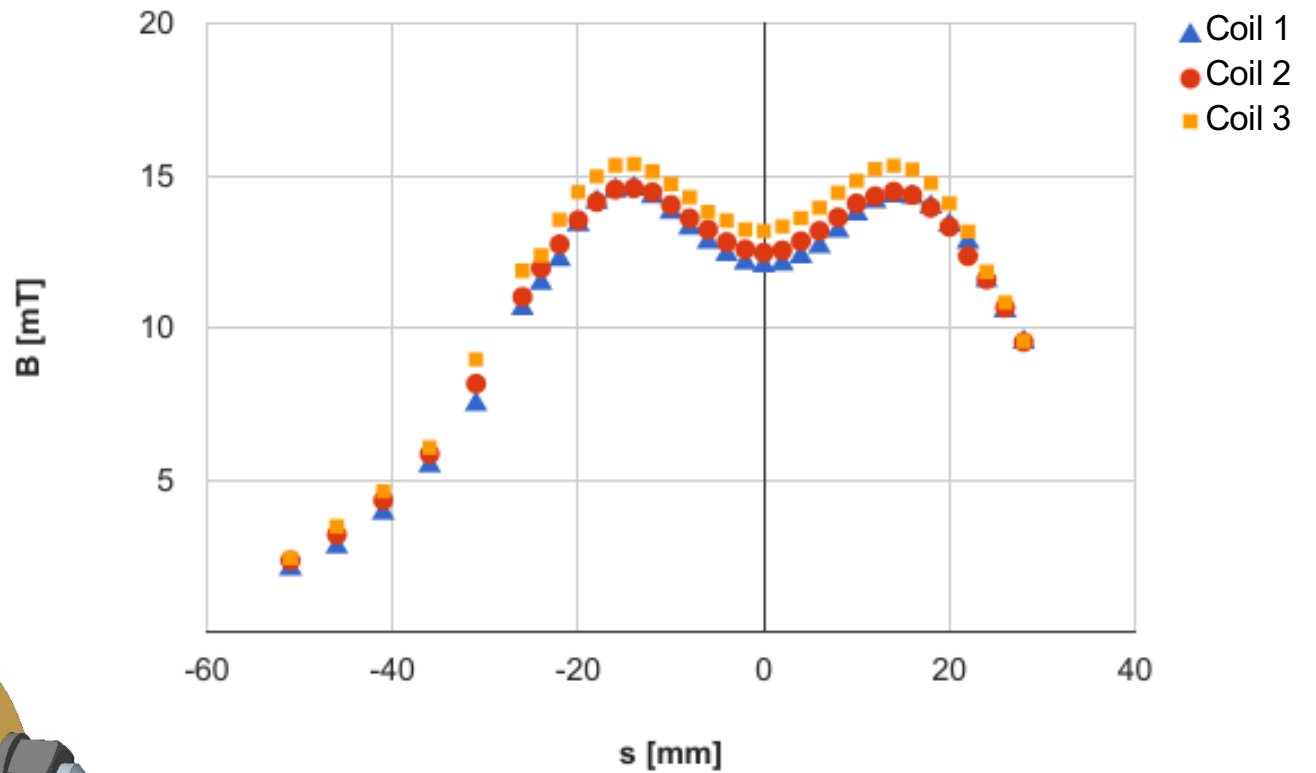
## Challenges:

- High field ( $\sim 20$  T) generation + optical access to gap
  - Strong attracting forces
  - High mechanical stress
- Vacuum compatible materials ( $\sim 10^{-5}$  mbar)
- Compact housing to prevent coil from outgassing
  - Flash over protection

# Magnet prototype



### Field measurement @ 20 A DC

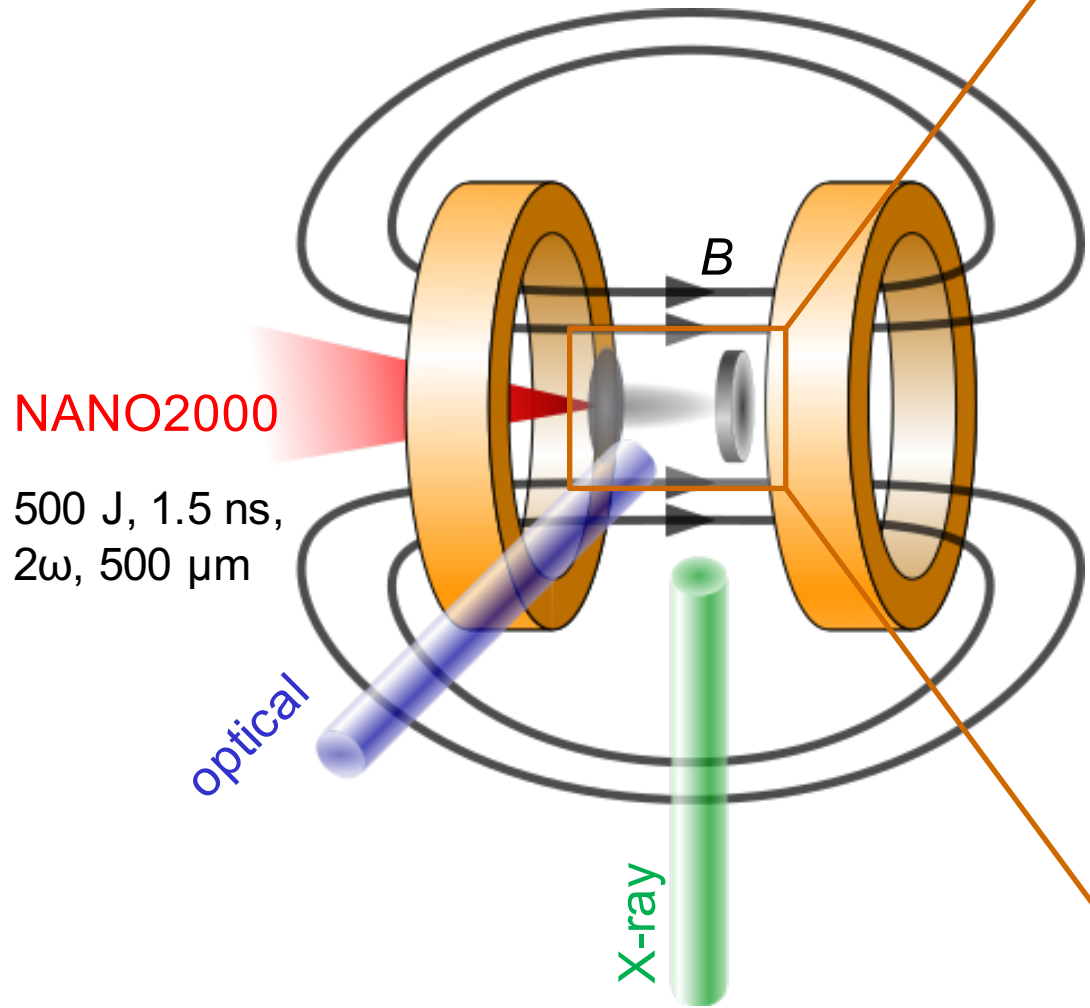


Successfully commissioned  
for 15 T operation

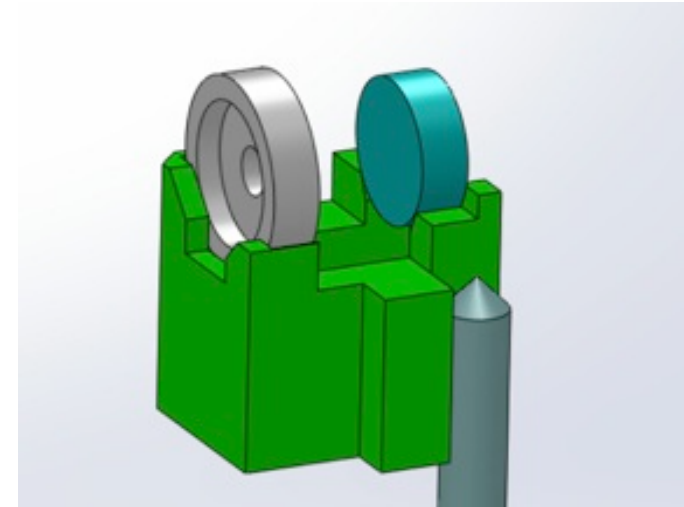
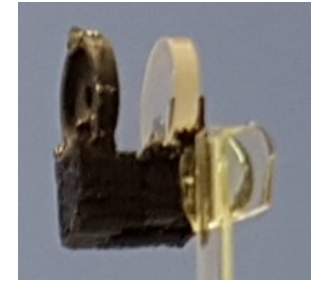




# Experimental setup @ LULI2000



## Target



### Laser targets:

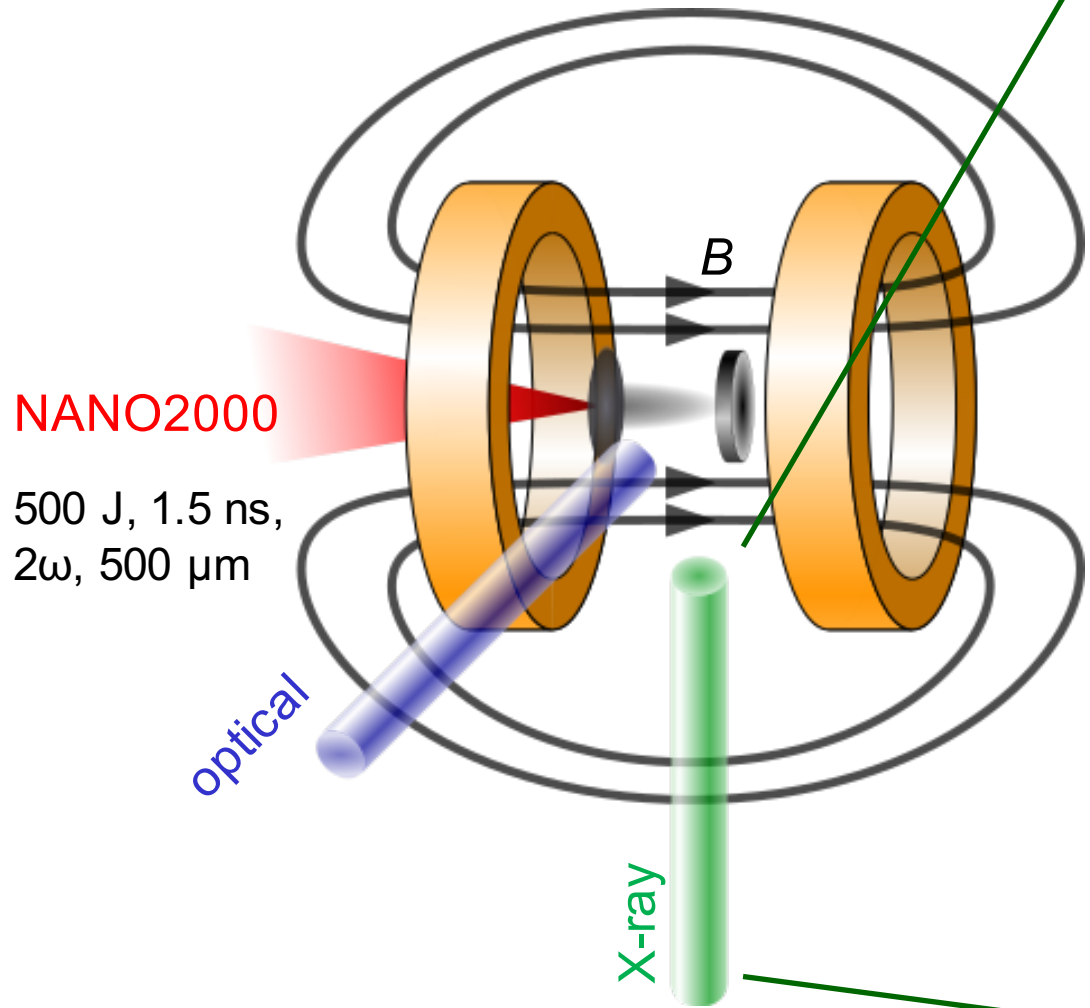
Al<sub>0.2 $\mu$ m</sub>-CH<sub>10 $\mu$ m</sub>-Au<sub>1.5 $\mu$ m</sub>-Ti<sub>6 $\mu$ m</sub>  
CH<sub>25 $\mu$ m</sub>-Sn<sub>6 $\mu$ m</sub>-CH<sub>25 $\mu$ m</sub>  
CH<sub>25 $\mu$ m</sub>-Sn<sub>6 $\mu$ m</sub>

### Obstacle:

GGG



# Experimental setup @ LULI2000



### X-ray generation

x-rays

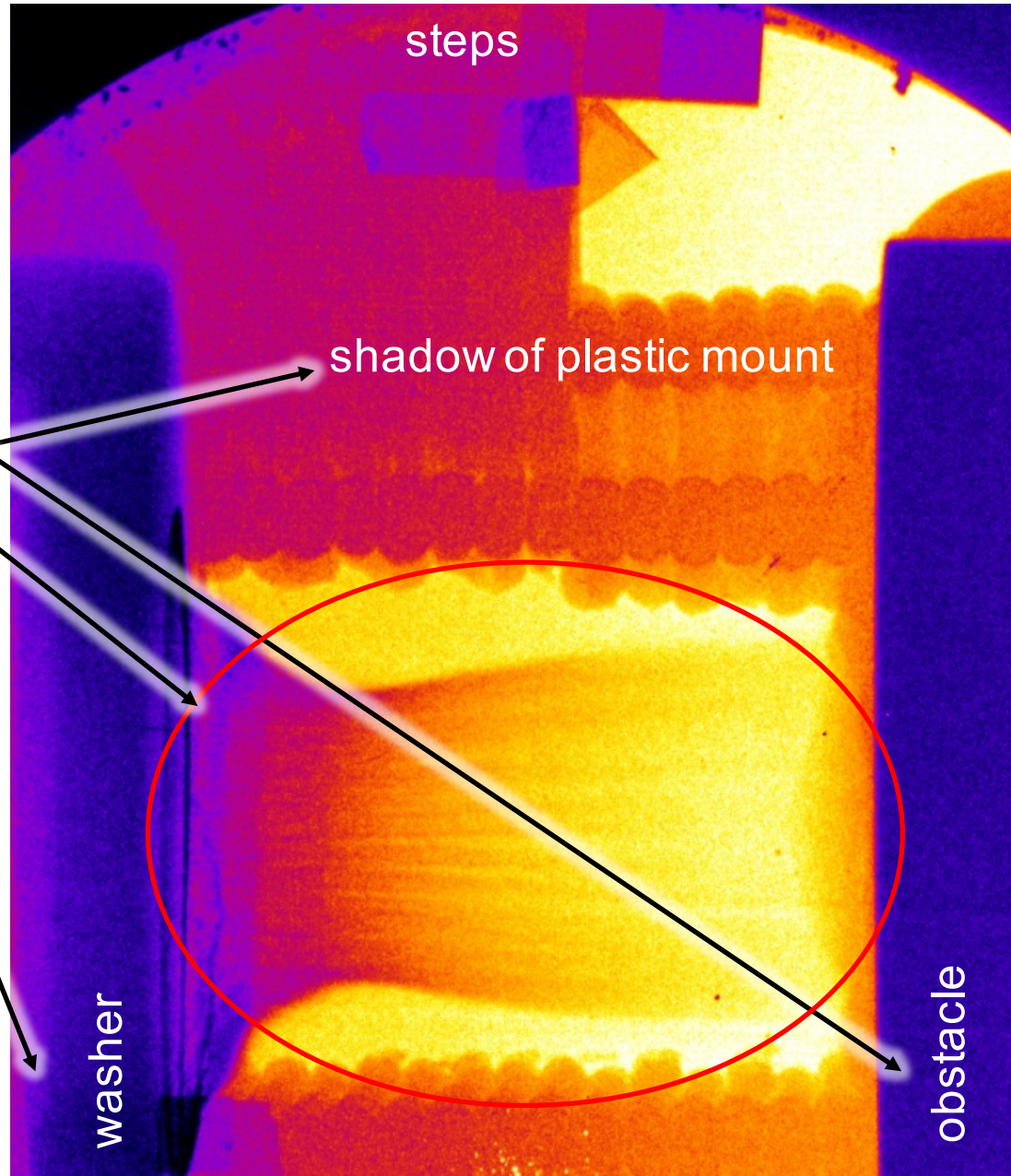
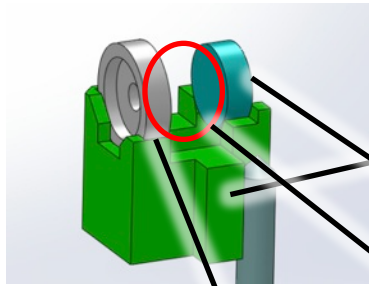
**PICO2000**  
80 J, 10 ps, 1 $\omega$

e<sup>-</sup>

source size  
25 $\mu$ m  
Ti wire

- X-rays detected by IP
- spectrally characterized
- ~ 20  $\times$  magnification
- 1.443  $\mu$ m/px
- 1000 lines/inch resolvable

# X-ray radiographs

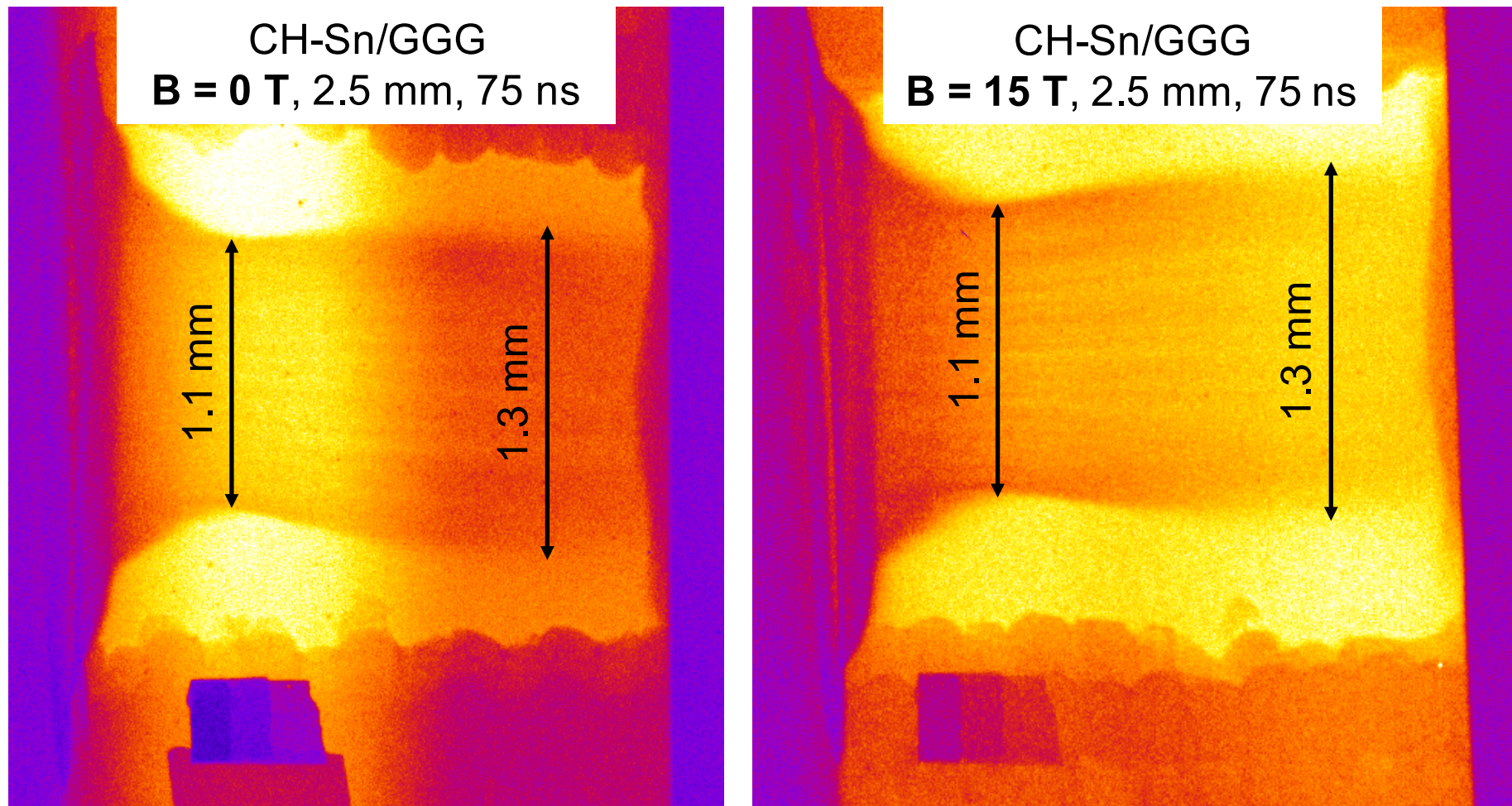


Magnetic field: 15 T,  
Distance W to O: 3 mm,  
Delay NANO-PICO: 120 ns





## w/o B-field vs. w/ B-field

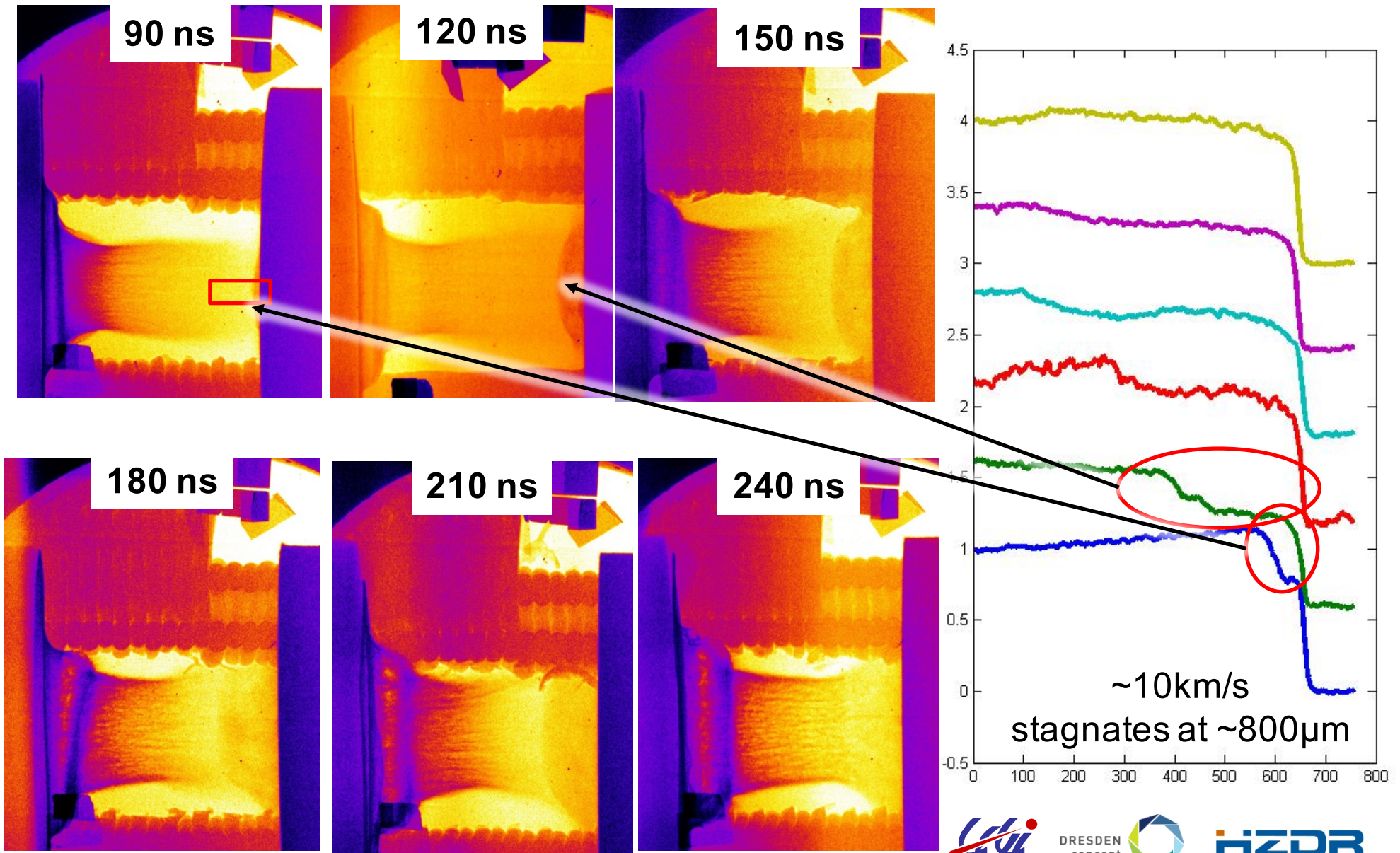


**Magnetic field has only little influence on high density region**

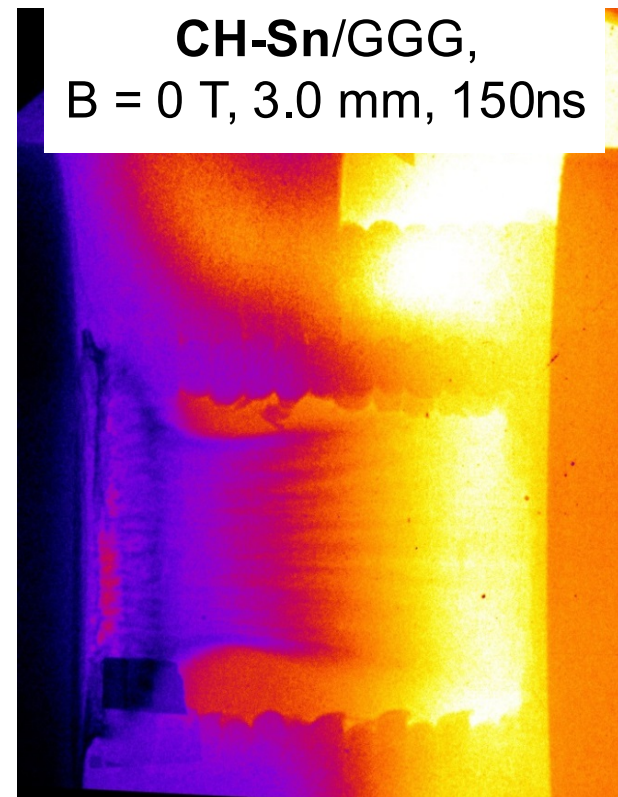
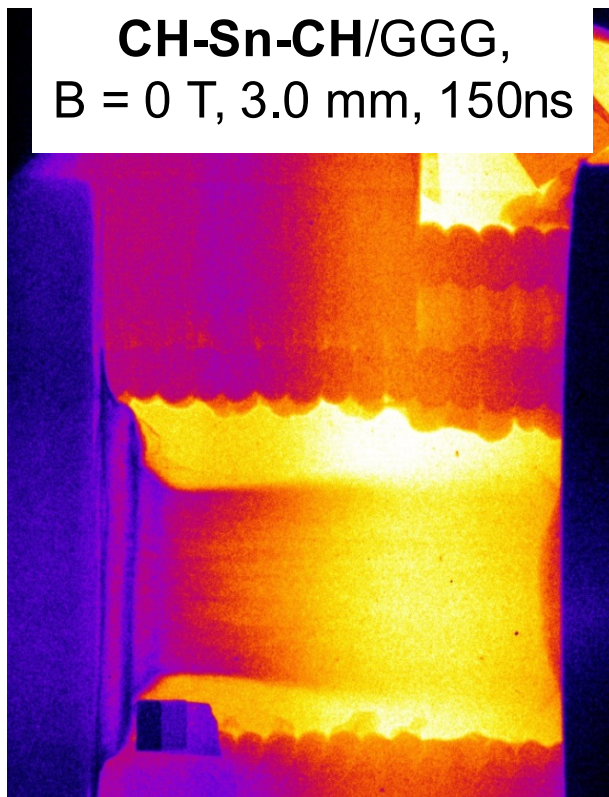




# Return shock @ different (long) delays, B = 15 T



## Different main targets

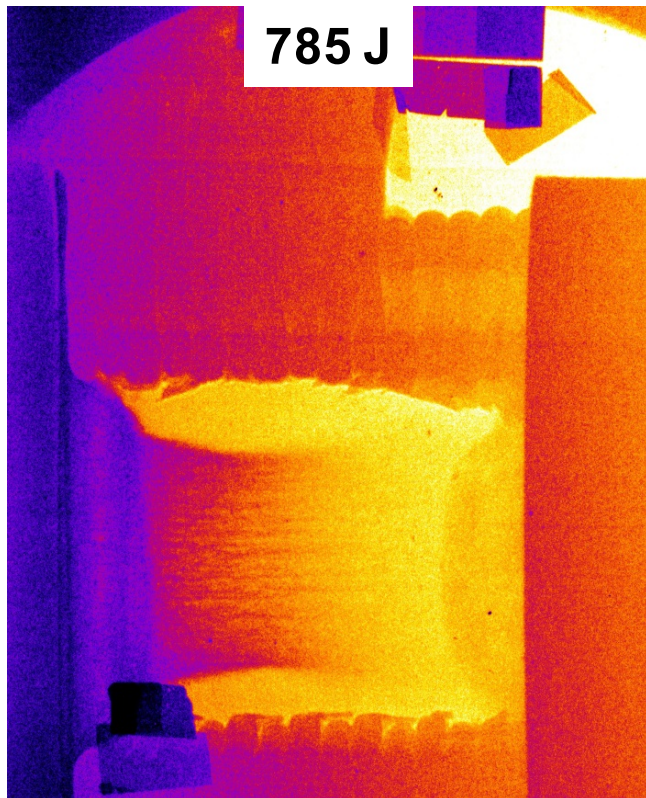


Analysis ongoing...

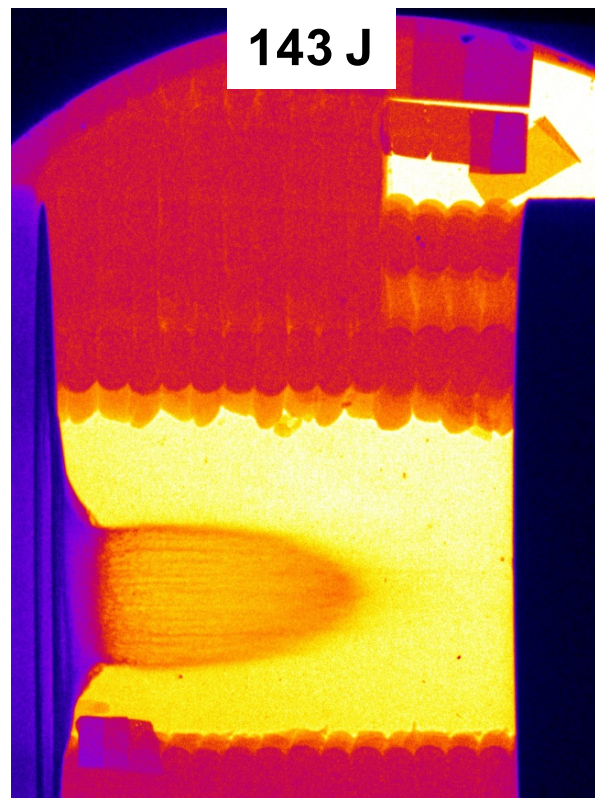




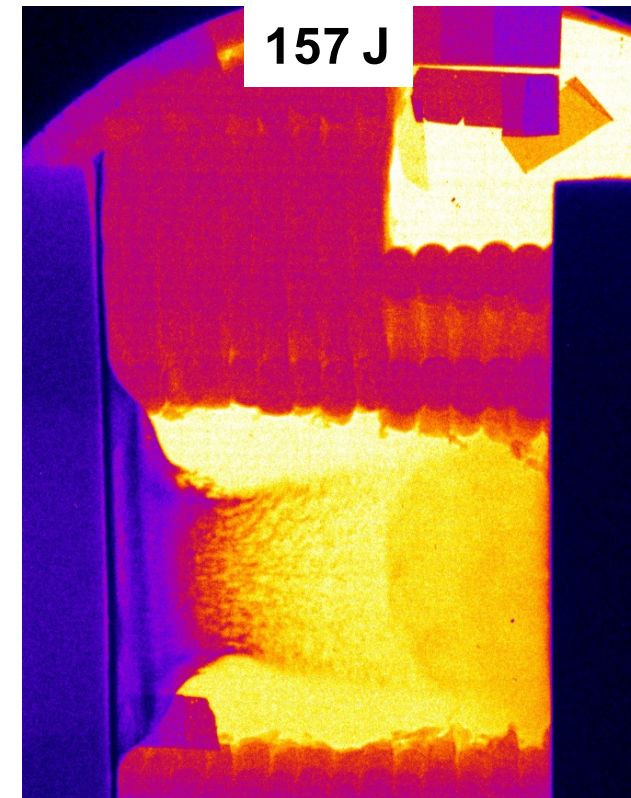
Reduced laser energy,  $B = 15 \text{ T}$



Al-CH-Au-Ti, 3 mm,  
150 ns



Al-CH-Au-Ti, 3 mm,  
145 ns



Al-CH-Au-Ti, 3 mm,  
395 ns

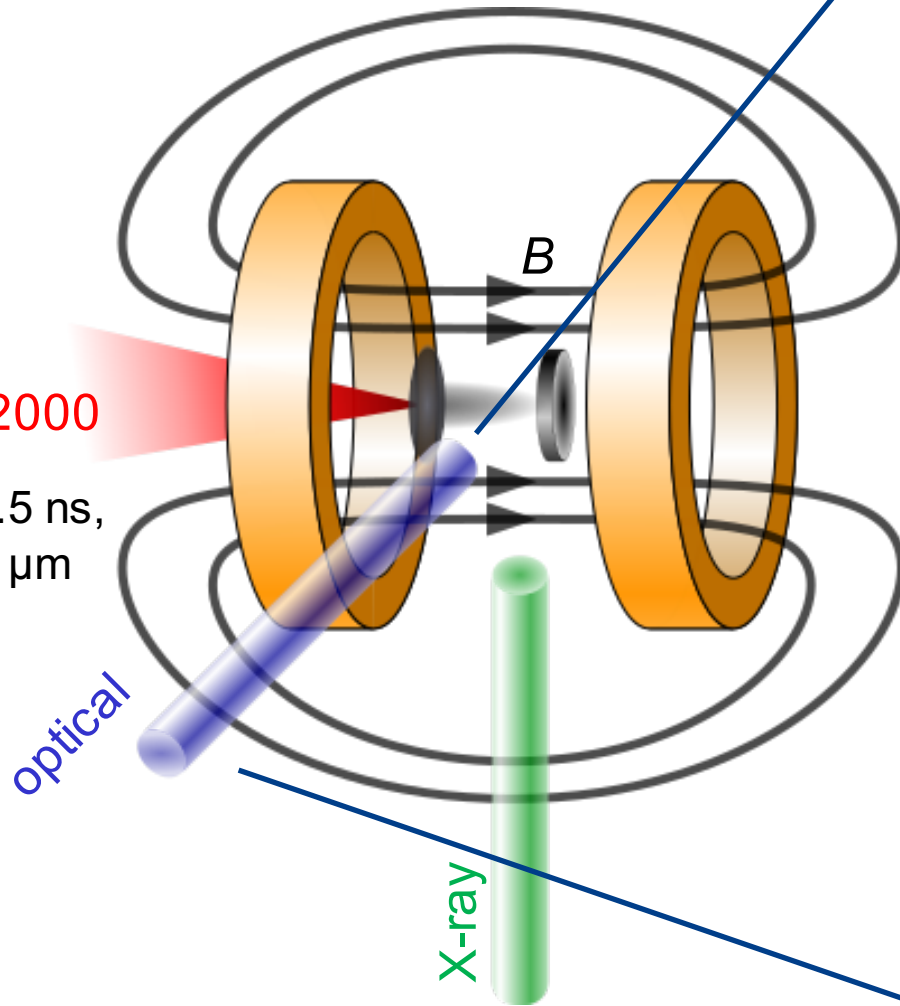
Analysis ongoing...



# Experimental setup @ LULI2000

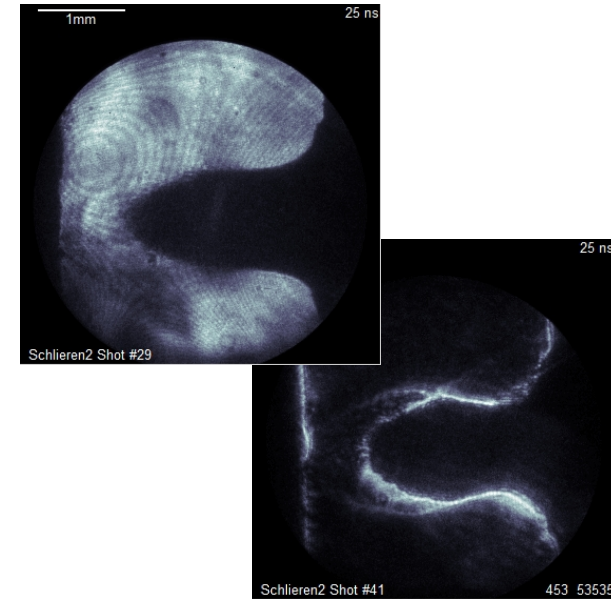
NANO2000

500 J, 1.5 ns,  
2 $\omega$ , 500  $\mu$ m



## transverse optical diagnostics

### shadowgraphy / schlieren

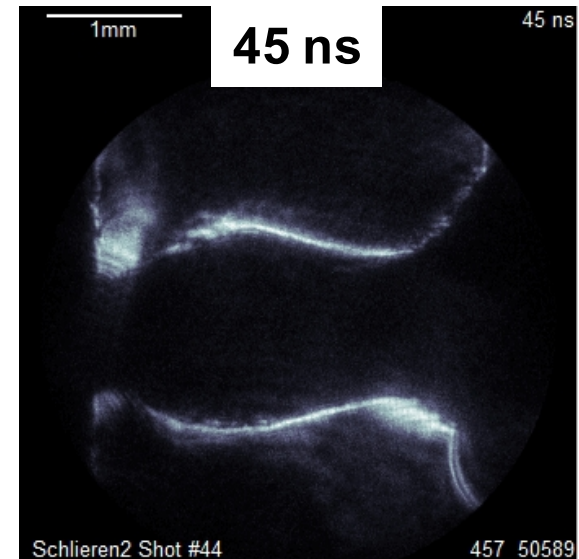
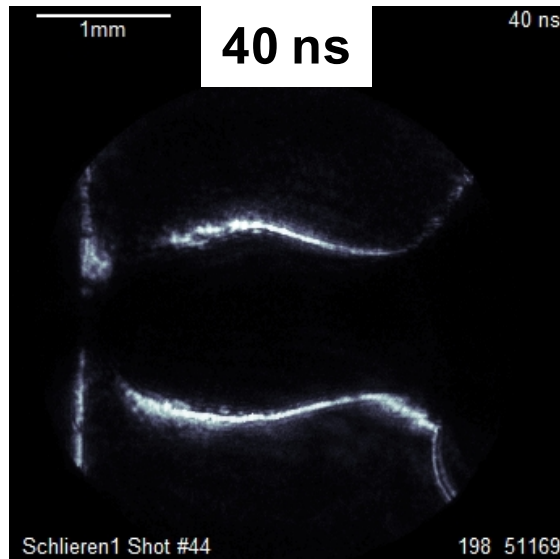
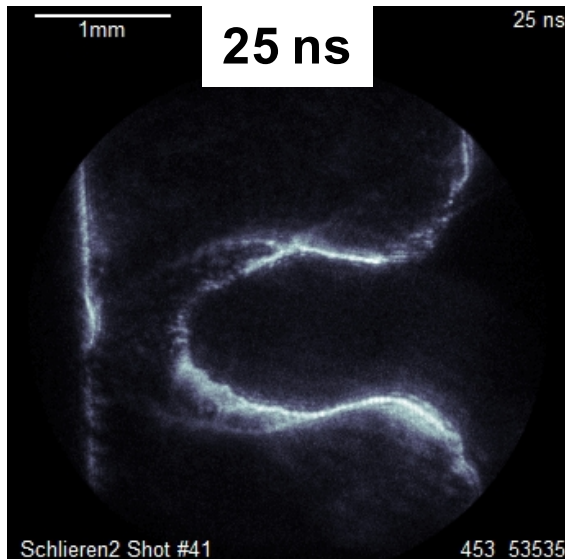
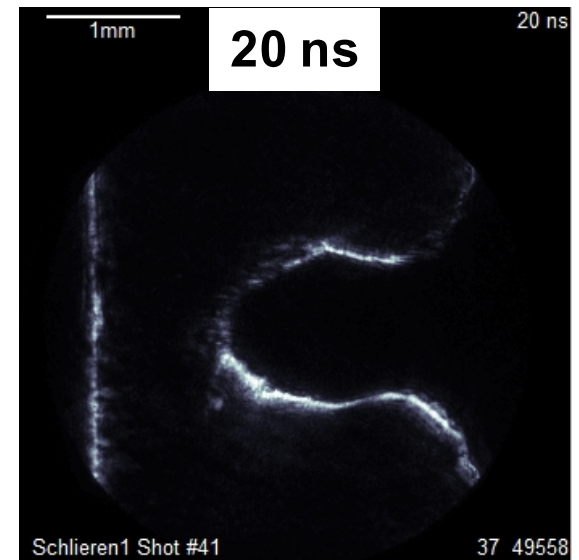
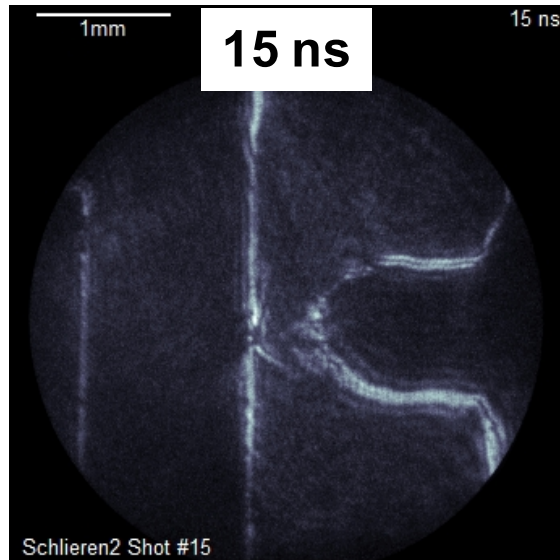
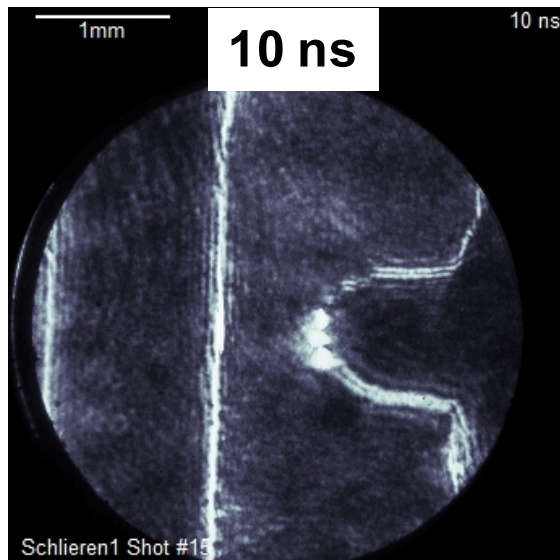


- 2  $\times$  schlieren imaging
- very sensitive to gradients in low density plasma
- difficult to quantitatively analyze





# Plasma expansion in Schlieren diagnostic



Collision of low density plasma with obstacle at ~40ns

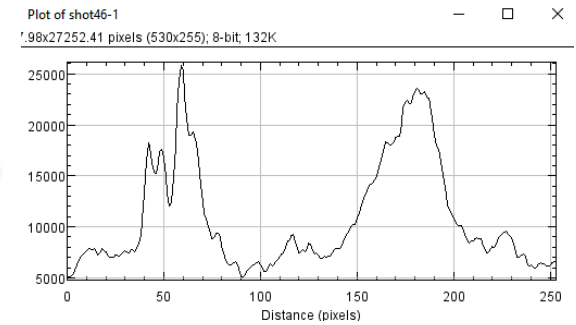
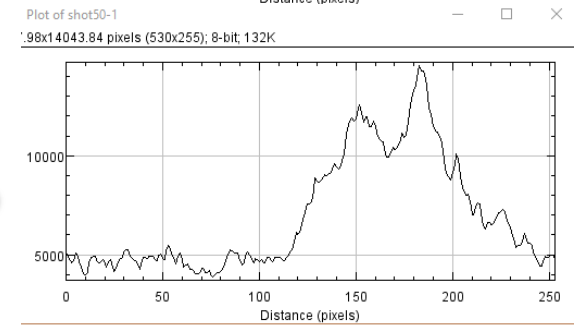
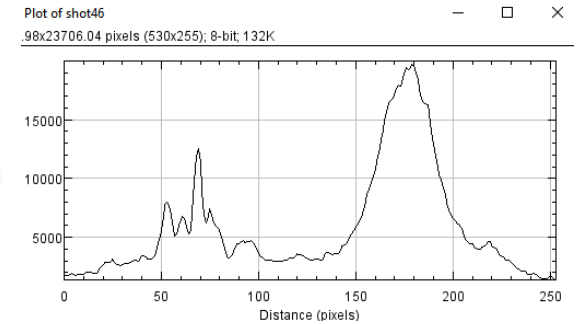
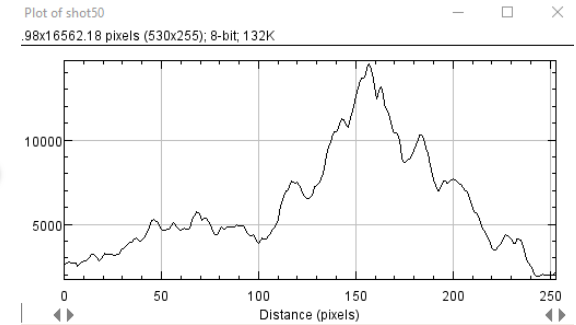
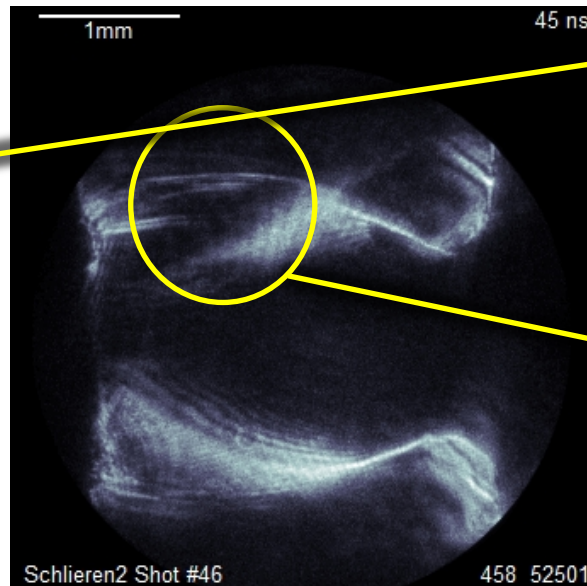
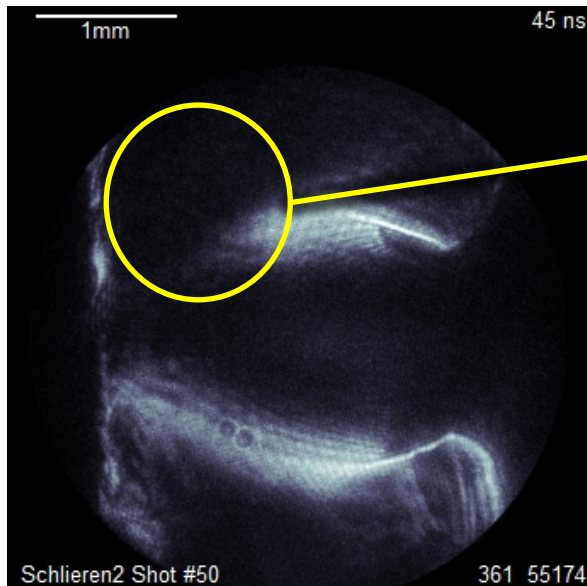
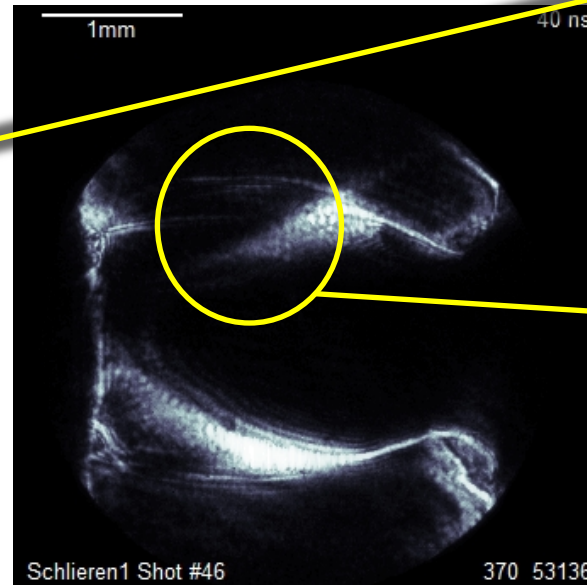
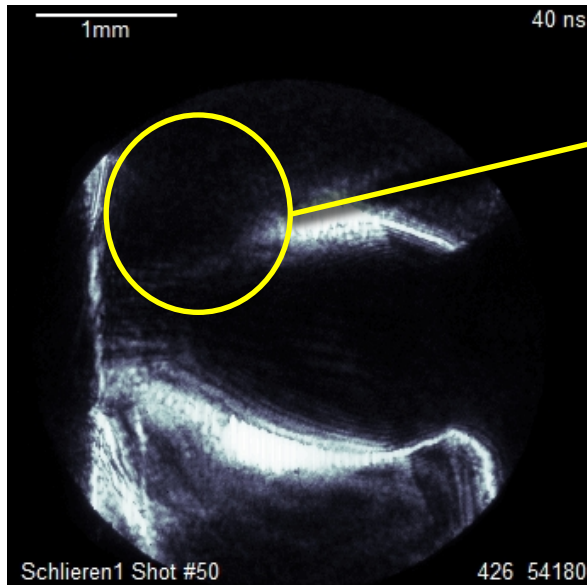




# w/o B-field vs. w/ B-field in schlieren

**B = 0 T**

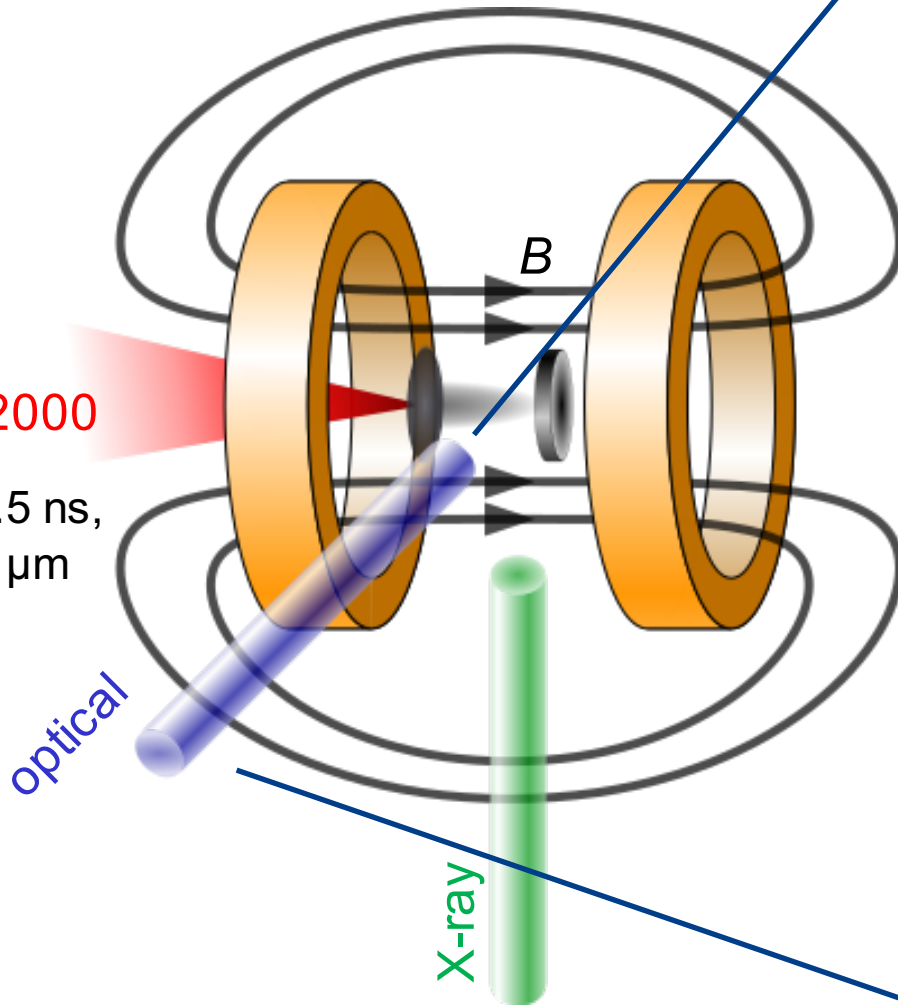
**B = 15 T**



# Experimental setup @ LULI2000

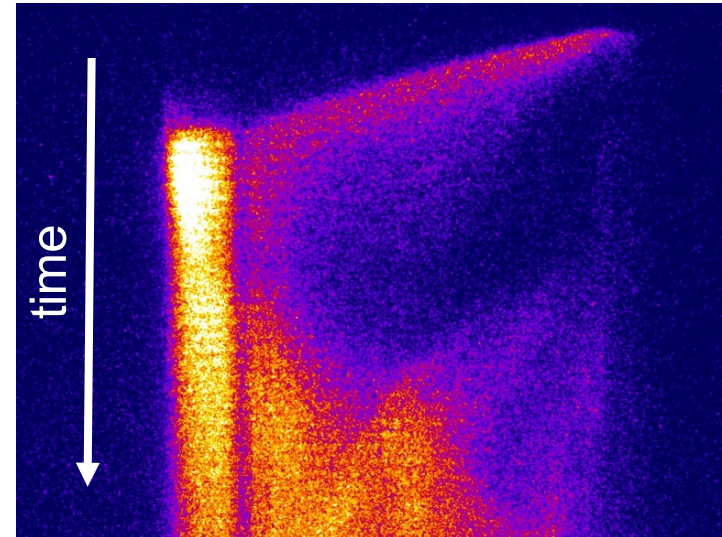
NANO2000

500 J, 1.5 ns,  
 $2\omega$ , 500  $\mu\text{m}$



transverse optical diagnostics

Streaked optical pyrometry  
(SOP)

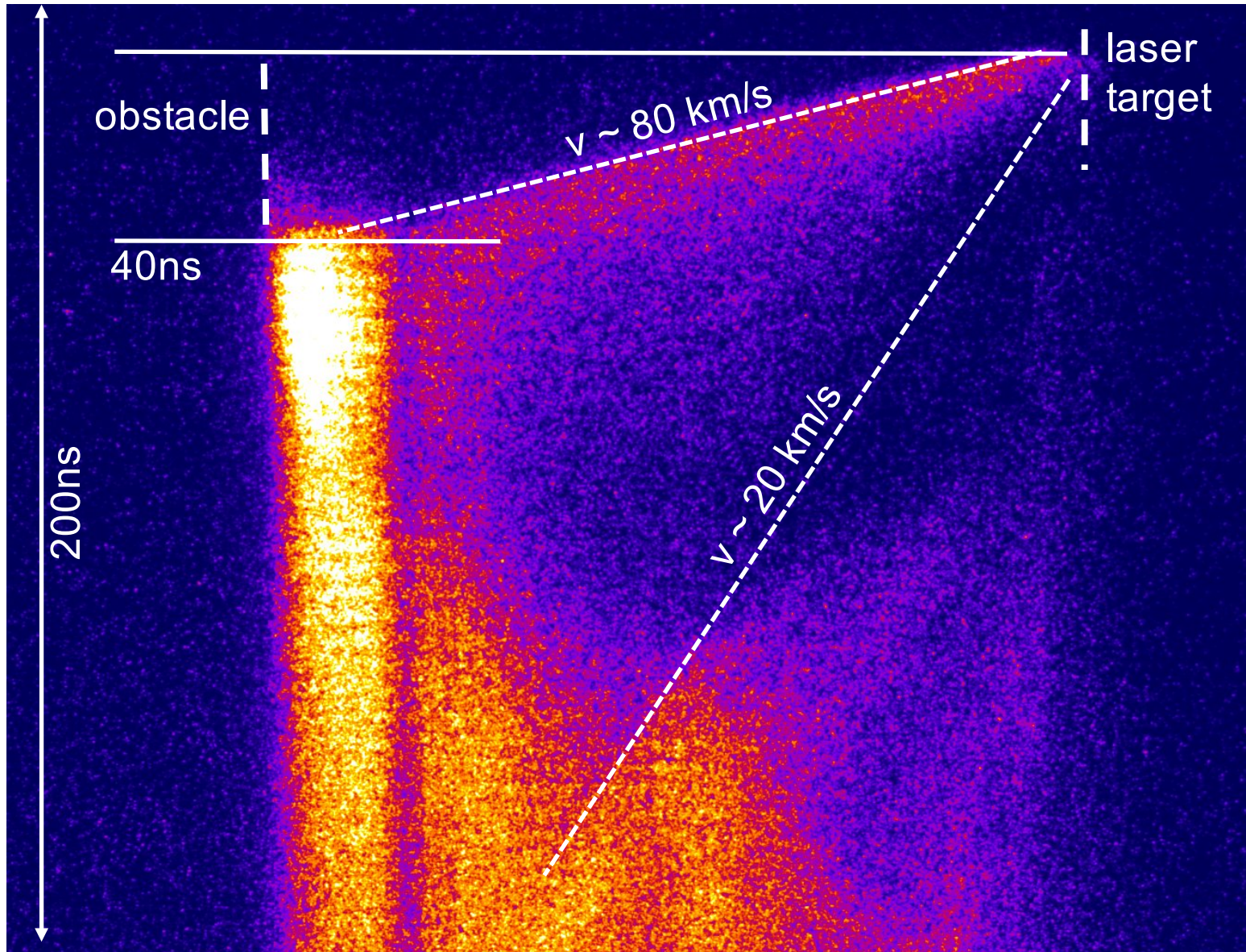


- collision time measurement
- determine plasma / reverse shock velocity





# Transverse SOP

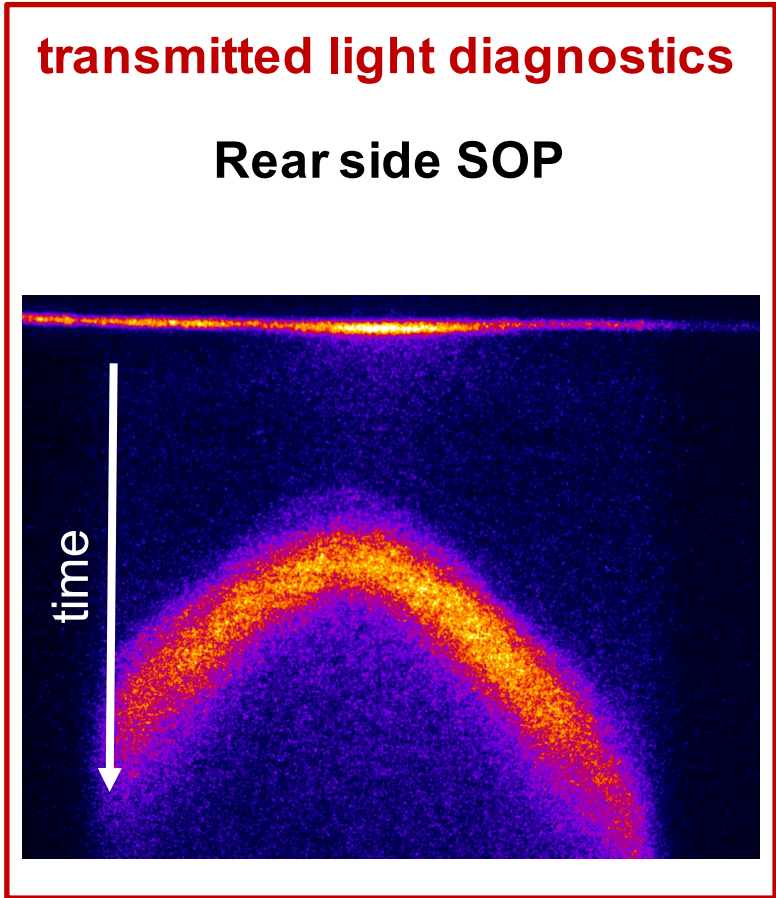
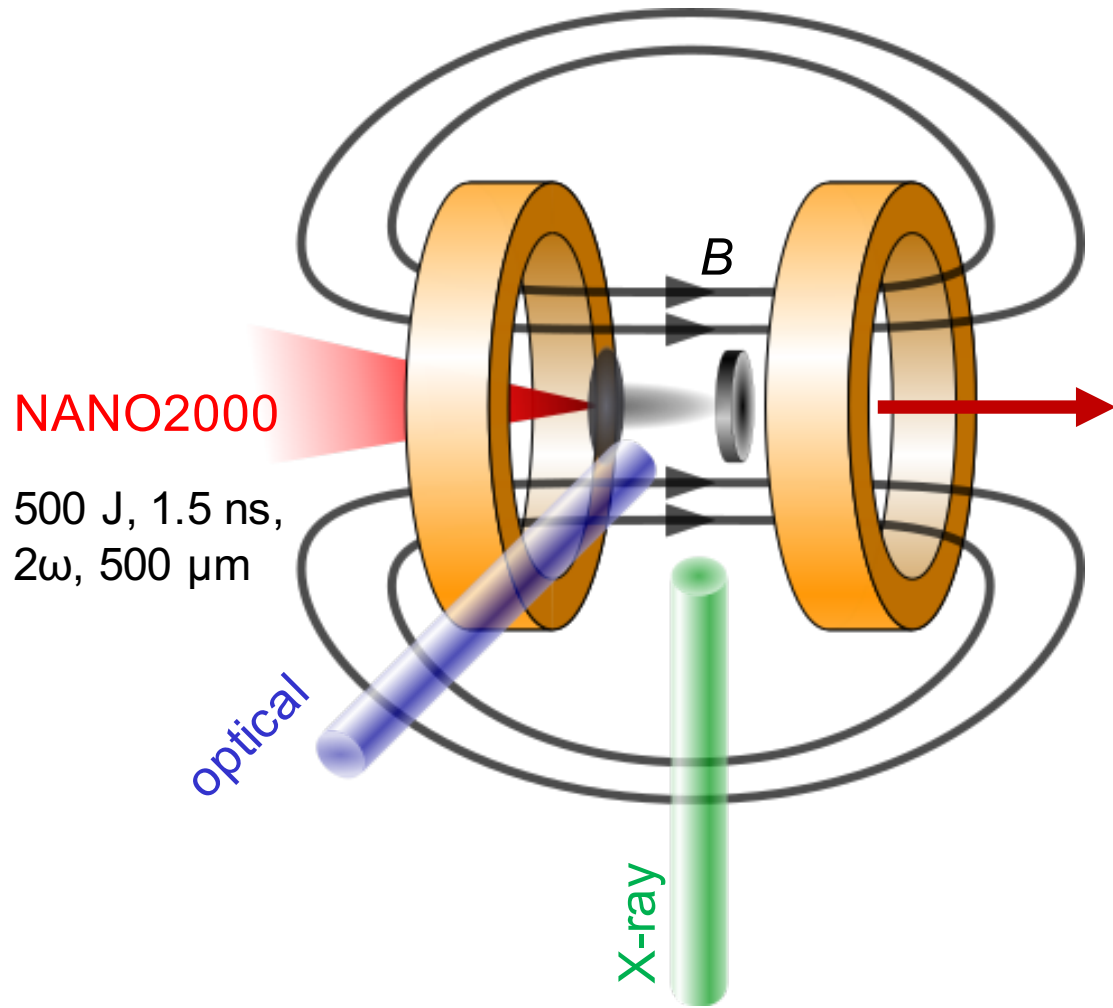


Collision time of 40 ns consistent with Schlieren

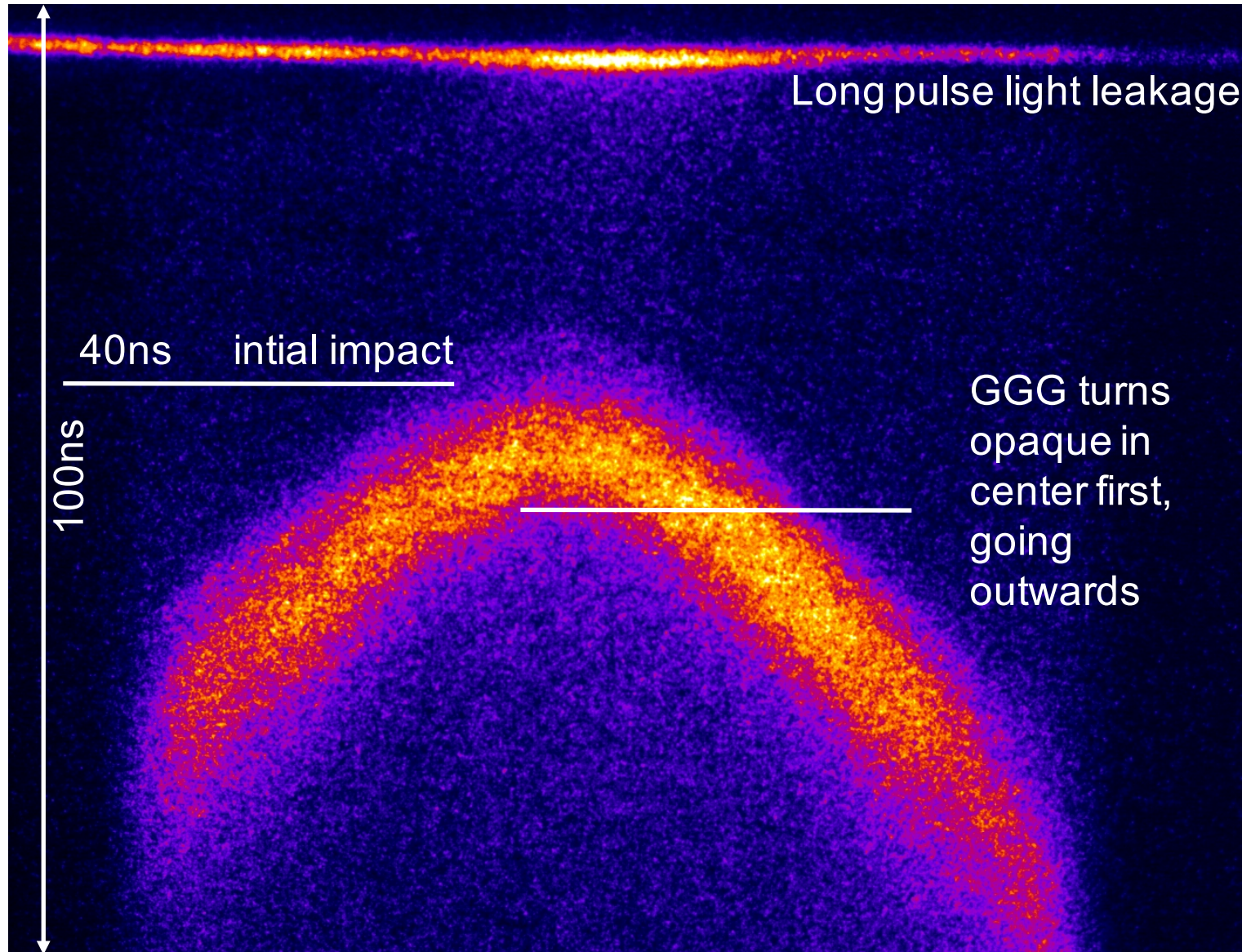




# Experimental setup @ LULI2000

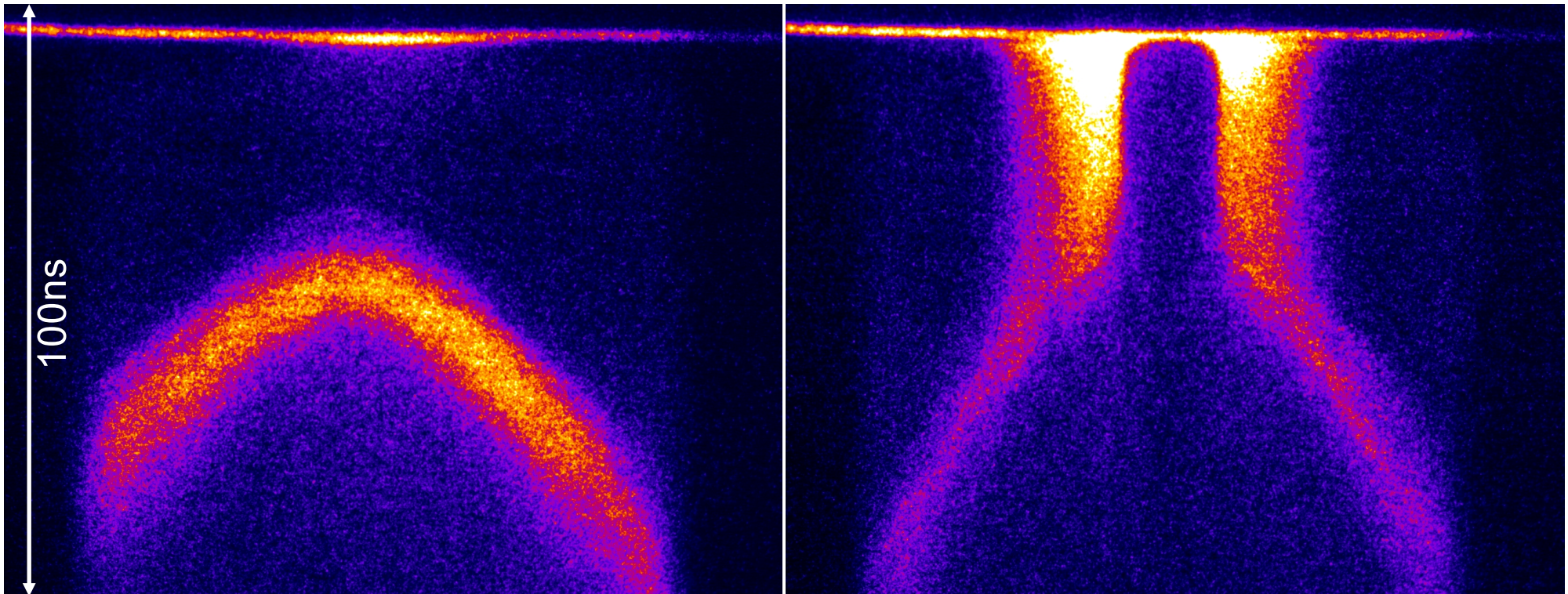


# Rear side SOP w/o B-field





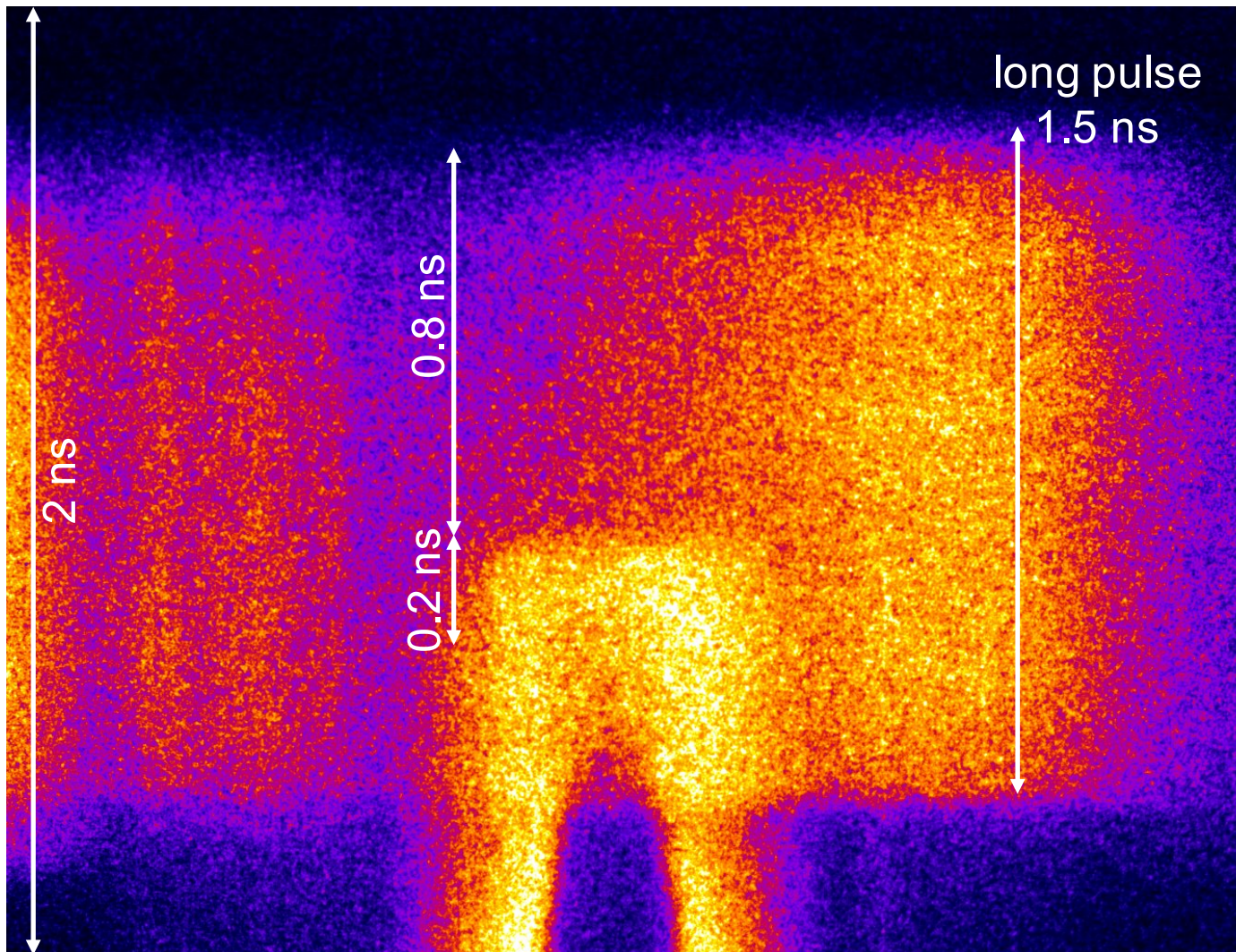
## Rear side SOP w/o B-field vs. w/ B-field



- GGG turns opaque very fast due to electrons locked to B-field
- Plasma outflow radiates stronger, but return shock collision shines weaker



## Rear side SOP w/ B-field – shorter time window



- shock breakout @~0.8ns
- GGG turns opaque @~1ns after long pulse onset
- quantitative analysis needed to determine if e<sup>-</sup> or other effect

## Summary

- Successful combination of pulsed coil with a longpulse/shortpulse laser facility
- generated laser-driven plasmaflows within the magnet
- diagnosed the plasmaflow both in the high density regime using laser-generated x-ray radiography and the low-density regime using various optical diagnostics
- generated and radiographed a return shock from the collision on hard obstacles

### Next steps:

- complete analysis of collected data
- next experiment: low-density, high-Z plasmaflows (foams), stronger magnet
- Lab-astro experiments at XFEL?

